ROSETTA

RPC-MAG SOFTWARE

RO-IGEP-TR-0007

Issue: 3 Revision: 1

January 20, 2010

DDS2PDS

User Manual

Ingo Richter

Institut für Geophysik und extraterrestrische Physik Technische Universität Braunschweig Mendelssohnstraße 3, 38106 Braunschweig Germany

ΡΟΓΕΤΤΑ	Document:	RO-IGEP-TR-0007
Π Ο Β Ε Ι Ι Α	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLΓ Technische Universität Braunschweig	Page:	Ι

Contents

1	Introduction	1
2	Usage of DDS2PDS	2
3	S/W-Routines of the DDS2PDS package 3.1 The DDS2PDS shell 3.2 DO_PDS - Initialization of the data processing 3.3 RAW2PDS - Transfer of binary data and conversion to PDS Data 3.4 GET_FTP_DATA 3.5 Conversion of the binary TM data: RAW2ASCII 3.6 Calibration of the converted data: GEN_CAL_DATA 3.7 GEN_LEVEL_KL_DATA - Generation of heater current corrected LEVEL_B data - (CLK) 3.8 GEN_LEVEL_KL_DATA - Generation of heater current corrected LEVEL_K data in ECLIPJ2000 coordinates - (CLL)	$\begin{array}{c} 3\\ 3\\ 3\\ 3\\ 3\\ 4\\ 5\\ 6\\ 7\\ 7\\ 8\\ 8\\ 8\\ 8\\ 8\\ 10\\ 10\\ 10\\ 11\\ 11\end{array}$
4	Output of DDS2PDS4.1Data product overview4.2Special handling of EQMTEST data	12 12 17
5	Directory and Filename Definiton5.1Software Directories5.2PDS Data Directories5.3Plot Directories5.4Temporary Working Directory5.5Input, Parameter and RAW2ASCII Directories5.6Output filenames5.6.1DATA files5.6.2PLOT files5.6.3Logging files	 18 18 20 20 21 21 21 23 24
6	The command file PDS.CMD and the control parameters of DDS2PDS6.1Syntax description	25 25

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
Institut für Geophysik u extraterr Physik	Revision:	1 January 20, 2010
IGEP Technische Universität Braunschweig	Page:	January 20, 2010 II

	6.2	Example of a Command File	43
7	File	format of RAW2ASCII files	46
	7.1	Input files	46
	7.2	Output files	46
		7.2.1 Housekeeping files	46
		7.2.2 Magnetic field files	47
8	File	format of PDS files	48
	8.1	Example: A generated PDS label file (*.LBL)	48
	8.2	Example: A generated PDS data file $(*.TAB)$	52
9	File	format of LATEXoutput file PLOTFILES.TEX	53
10	Anc	illary IDL Routines	55
	10.1	Generation of a Temperature Model	55
	10.2	Generation of REACTION WHEEL Plots	60
		10.2.1 Reaction Wheel data from DDS-Server	60
		10.2.2 Reaction Wheel data from IC-Server	61
		10.2.3 Processing of the Reaction Wheel data	61
	10.3	Generation of Solar Array (SA) and Hight Gain Antenna (HGA) Plots	62
	10.4	Generation of overview plots and overview about available data	63
	10.5	Generation of a Operations Logbook	64
11	Arcl	hive Generation Procedure	65
12	Note	es	67
	12.1	Plotting Parameters	68

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 1

1 Introduction

This document describes the data processing S/W *DDS2PDS* which is used for processing and archiving of the ROSETTA RPCMAG data. The S/W has been developed to provide a quasi–automatic data processing chain from the binary magnetic field data on the ESA Data distribution system (DDS) to the data archive at the Imperial College (IC) server which is compliant to the Planetary Data System (PDS) rules.

The DDS delivers binary files with magnetic field raw data of the RPCMAG Inboard (IB) and Outboard (OB) sensor, Sensor temperatures of the IB and OB sensor and additionally RPCMAG Housekeeping (HK) data. The magnetic field data which is delivered as 20 bit ADC two's complement data in instrument coordinates have to be converted in the following way. They have to be rotated into s/c-coordinates, the ground calibration parameters have to be applied, and a temperature correction has to be performed. The result of this procedure will be calibrated data. On a higher level we will rotate these data in a convenient celestial body frame (e.g. ECLIPJ2000, CSO,...) and average these data to a convenient rate (e.g. 1s mean). The s/w can read Lander heater corrected data and feed them into the standard data processing algorithms. The s/c generated offsets and the structures arising from the s/c noise are NOT removed in these data. The elimination of these effects is under development and will lead to derived data products (TBD)e.g PCA data.

Finally plots can be generated and the data will be written to PDS compliant Label *.LBL and Table files *.TAB. These files can automatically be saved on the IGEP ftp–server and will be transferred automatically to the Imperial College Data Server (Pegasus) by the means of a cron job.

The IDL S/W package DDS2PDS consists of several routines for different purposes which will be extensively described in this document. The main features of the S/W are:

- Copying TM raw data from the IGEP ftp-server to the local analysis PC
- Converting /Decoding these binary data to ASCII data. This is done by calling the Matlab S/W *RAW2ASCII* from the IDL program.
- Reading attitude and orbit files (*.ROS) from the local geometry directory ROSETTA/DATA/GEOMETRY/.
- Calling the ESTEC OASWLIB S/W to generate desired attitude and orbit vectors.
- Generating PDS files from these ASCII raw data (Routine: *GEN_CAL_DATA*). Here edited raw data, calibrated data in different levels and averaged data can be produced.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 2

- Generating plots (Routine: *GEN_PLOTS* in the file *PLOT_DAY.PRO*)
- Copying the produced PDS *. TAB and *. LBL files to the IGEP ftp-server.
- Generating log–files and distributing them in automatically generated mails.

The package was developed under IDL running on a WINDOWS–XP platform.

2 Usage of *DDS2PDS*

The program has to be started via the batch file @dds2pds in the directory /Rosetta/rpcPds/rpcMag/SOFTWARE/PCWIN/IDL.

The S/W can be executed in two modes:

- Manual Mode: All parameters for the data processing have to be typed in during runtime after being prompted.
- Batch Mode: *DDS2PDS* reads the command file *PDS.CMD* with all needed parameters. This mode is recommended for the standard analysis.

The parameters of the manual mode are all available in the Batch mode. However, the Batch mode has a lot more capabilities. All parameters are described in section 6.

ΡΟς ΕΤΤΛ	Document:	RO-IGEP-TR-0007
ΠΟΒΕΙΙΑ	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
Ι GΓΓ Technische Universität Braunschweig	Page:	3

3 S/W-Routines of the DDS2PDS package

The following list gives a summary of the complex data transfer, calibration and archiving process. Most of the steps can be controlled by control variables. These variables can be read manually or from the *PDS.CMD* file. Without restriction of generality the following list of routines describes the process with all variables set to the activated status.

3.1 The DDS2PDS shell

This shell initializes the global variables, compiles all routines and calls the DO_PDS routine finally.

3.2 DO_PDS – Initialization of the data processing

- 1. Initialization of data process control variables. This is done either by manual setting or by reading the command file *PDS.CMD* One of the most important parameters is the DATE (variable name: DAY) of the data to be processed. The software processing is organized in that way, that always data of COMPLETE days will be processed.
- Inquiry on the Mission Phase to be processed.
 The system has to know the mission phase to set the variables in the *.LBL files to the right values. In former version this had to be done manually, now the mission phase will automatically be identified by the date of the day to be processed.
- 3. The internal procedure *START_PROCESSING* is called.
 - The conversion process starts by calling RAW2PDS
 - If resampled LEVEL_E_DATA ... LEVEL_J_DATA shall be generated, the procedure *GEN_LEVEL_E_DATA* ... *GEN_LEVEL_J_DATA* will be executed.
 - If plots shall be generated, the plot generation routine *GENERATE_PLOTS* will be called.

3.3 RAW2PDS – Transfer of binary data and conversion to PDS Data

- 1. Initialize environment variables for directories and standard filenames.
- 2. Then the data transfer process starts by calling GET_FTP_DATA .

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 4

- 3. If the ASCII data is available on the local computer the calibration procedure GEN_CAL_DATA is called.
- 4. After the successful generation of ASCII data on the local computer these data can be transferred to the IGEP ftp–server by the means of the procedure *WRITE_PDS2FTP*.

3.4 *GET_FTP_DATA*

- 1. Initialize environment variables for directories and standard filenames. Especially the external batch file names are defined here, as the major part of the data transfer is done by external routines and programs.
- 2. Delete the
 - input directory for *RAW2ASCII* (rawdir) : D:/rosetta/data/raw_files
 - local directory (localdir): D:/rosetta/data/ftplocal
 - *RAW2ASCII*–Output directory (ascdir) : D:/rosetta/data/asc_files
 - temporary directory (tempdir): D:/rosetta/rpcPds/rpcMag/tmp
- 3. Generate the script file to get the needed directory listing from the ftp–server (procedure: gen_dir_scr)
- 4. Transfer the directory listing of all available data from the ftp-server (procedure: get_ftp_dir)
- 5. Now this listing has to be filtered and the right files have to be transferred:
 - Select MAGNETIC FIELD data: Standard search path for the magnetic field data on the IGEP-ftp server is: /home/mag/rosetta/rpcData/archivex/tlm/yyyy/mm/dd/ rpc<yymmdd>dds000_tlm_all____<start_time>_<end_time>.tm All MAG data filenames of the desired day are filtered out of the list (procedure: *read_mag_dir*). If there are any files, these extracted names will be written to a list (procedure: *gen_mag_copy_file*) and the datafiles of this list will be transferred from the ftp-server to the local machine (procedure: *get_mag_data*) to the *localdir* directory.
- 6. After this preprocessing the magnetic field data file is copied to the rawdir directory as (MAGyyyymmdd.ALL)
- 7. Also the auxiliary files for the Reaction wheel disturbance correction *schk7.txt* and the Instrument status file *piu_hk.txt* are transferred from the IGEP-ftp server */home/mag/rosetta/rpcData/archivex/tlm/yyyy/mm/dd* to the local directories

ROSETTA	Document: Issue:	RO-IGEP-TR-0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 5

- d:\rosetta\data\reaction_wheels\ascii_data\schk7_yyyy-mm-dd.txt
- d:\rosetta\data\reaction_wheels\ascii_data\ piu_hk\piu_hk_yyyy-mm-dd.txt
- 8. Now the raw data conversion S/W (*raw2ascii*) is called directly from IDL as a batch process. Input of (*raw2ascii*) is the binary (*MAGyyyymmdd.ALL*) file which contains all HK data and every magnetic field data of the desired day data in every mode and from every sensor. The output of this conversion S/W are separate ASCII files: one file for every mode and sensor. All output files (extension *.*ASC*, ASCII) are written to the (*ascdir*). The exact description of (*raw2ascii*) is given in section 3.5.
- 9. It turned out that the content of the (*raw2ascii*) outputfiles is unsorted. This is due to the flexibility of the RPCMAG instrument, which is able to set either the OB or the IB to the primary sensor. So the outputfiles have to be sorted (this is done by the tool "gawk" in conjunction with "sort"). Additionally double entries have to be removed (achieved with the tool "uniq"). The output files of these additional process have the extension (*.ASS, ASCII sorted) and can be found in the *assdir* which is set to *D:/rosetta/data/ass_files*

REMARK:

In case that the raw data files shall not be downloaded from the ftp server, but directly be loaded from the local machine during the GET_DATA process, the following procedure has to be applied:

- copy *.tlm file as MAGyyyymmdd.all to data\raw_files
- copy the *schk7.txt* and *piu_hk.txt* files to the right local filenames
 - d:\rosetta\data\reaction_wheels\ascii_data\schk7_yyyy_mm_dd.txt d:\rosetta\data\reaction_wheels\ascii_data\piu_hk\piu_hk_yyyy_mm_dd.txt
- After starting *DDS2PDS* advise the s/w to load the files locally by typing an "L" at the related prompt.

3.5 Conversion of the binary TM data: RAW2ASCII

Binary TM data will be read and converted to ASCII by RAW2ASCII.

Program details:

 $\bullet~RAW2ASCII$ is developed in MATLAB under Windows by Hans Eichelberger, IWF, GRAZ

ROSETTA	Document: Issue:	RO-IGEP-TR-0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 6

- *RAW2ASCII* acts as i/f between the binary raw data transmitted by the DDS or the EGSE or the IC–ftp-server (standard processing) server and the scientific usable data.
- The program converts binary raw data into ASCII data and adds the necessary time information (UTC and s/c-clock) for the subsequent scientific analysis. Bad vectors are marked. All written ASCII files get a header starting with #. The exact file format is described in section 7.
- It reads
 - Magnetic field raw data in all modes (SID1 SID6)
 - Temperature data (IB/OB)
 - Housekeeping data (HK)
- The program can be executed via a batch job to guarantee a more ore less automatic data generation/conversion process.

3.6 Calibration of the converted data: GEN_CAL_DATA

- The converted ASCII data will be merged with auxiliary data (Attitude & Positions) and processed with *GEN_CAL_DATA* to obtain scientific usable data in PDS format.
- This IDL routine acts as I/F between the ASCII raw data converted by *RAW2ASCII* and the PDS System.
- *GEN_CAL_DATA* reads (files can be read from a list for automatic data generation)
 - Magnetic field ASCII raw data: *RPCMAGyymmddThhmm_RAW*<sensor>_<MODE>.ASS
 - Auxiliary data Attitude: ATNR*.ROS
 - Auxiliary data Position: OR*.ROS
 - Housekeeping data: *RPCMAGyymmddThhmm_RAW_HK.ASS*
 - GND Calibration files: *RPCMAG_GND_CALIB_FSDPU_FM*<sensor>.TXT
 - Inflight Calibration files: $RPCMAG_{-} < id > CALIB_{-} < sensor > .TXT$ These two files are re-written day by day with the actual field values from the start and the end vector of the actual file.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 7

- Offset correction files: CLA_OFFSETS_<sensor>.TXT
- Boom alignment file: *RPCMAG_SC_ALIGN.TXT*
- Functions of *GEN_CAL_DATA*:
 - 1. apply temperature dependent ground calibration results.
 - 2. apply temperature dependent inflight calibration results to get B-field in unit coordinates.
 - 3. turn B-field from instrument to s/c coordinates
 - 4. apply attitude data to get B-field in EME2000 or ECLIPJ2000 frame (or a similar one)
 - 5. apply filters, spike detectors,.... data processing routines to get "scientific usable" magnetic field data in ASCII time series.
- *GEN_CAL_DATA* writes PDS compliant calibrated data files and labels on different stages (*.TAB, *.LBL).

3.7 *GEN_LEVEL_KL_DATA* - Generation of heater current corrected LEVEL_B data - (*CLK*)

This routine is used to convert external heater corrected data (*.TAB) in LEVEL_B format into LEVEL_K data. The format is the same as the LEVEL_B format. The input data are taken from the D:\ROSETTA\DATA\HEATER directory. Output of the routine are *.LBL and *.TAB files.

3.8 GEN_LEVEL_KL_DATA – Generation of heater current corrected LEVEL_K data in ECLIPJ2000 coordinates – (CLL)

This routine is used to rotate the heater corrected data LEVEL_K data from s/c-coordinates to ECLIPJ2000 coordinates. The format is the same as the LEVEL_C format. As input data LEVEL_K data are taken. Output of the routine are *.LBL and *.TAB files.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 8

3.9 GEN_LEVEL_EF_DATA – Generation of averaged and merged LEVEL_A data – (CLE)

This routine is used to merge data of all modes collected on one day, calculate averaged data from these input files and save it as PDS data. As input the calibrated LEVEL_A data in instrument coordinates are taken. The average period has to be stated by the means of the variable RESAMPLED_AVERAGE= \sqcup . Output of the routine are *.*LBL* and *.*TAB* files.

3.10 GEN_LEVEL_FF_DATA – Generation of averaged and merged LEVEL_B or LEVEL_K data – (CLF)

This routine is used to merge data of all modes collected on one day, calculate averaged data from this input files and save it as PDS data. As input the LEVEL_B or LEVEL_K data are taken. The average period has to be stated by the means of the variable RESAMPLED_AVERAGE= \Box . The variable LEVEL_F_SOURCE has to be set. If not stated LEVEL_B will be taken as input. Output of the routine are *.LBL and *.TAB files.

3.11 $GEN_LEVEL_G_DATA$ – Generation of averaged and merged LEVEL_C or LEVEL_L data – (*CLG*)

This routine is used to merge data of all modes collected on one day, calculate averaged data from this input files and save it as PDS data. As input the LEVEL_C or LEVEL_L data are taken. The average period has to be stated by the means of the variable RESAMPLED_AVERAGE= \Box . The variable LEVEL_G_SOURCE has to be set. If not stated LEVEL_C will be taken as input.Output of the routine are *.LBL and *.TAB files.

3.12 GEN_LEVEL_H_DATA – Generation of reaction wheel corrected LEVEL_C or LEVEL_L data – (CLH)

It turned out that that rotation frequency (or better: the folded frequency with respect to the actual sampling frequency according the the Nyquist Theorem) of ROSETTA's reaction wheels (RW)can be seen in the magnetic field data. Therefore, a RW-correction algorithm has been developed and implemented to this s/w. The program reads the actual reaction wheel frequency file containing the time and the frequency of the 4 wheels. These data are used to eliminate the traces of the wheels in the spectra and to generate corrected time series as PDS LEVEL_H data. As input the LEVEL_C or LEVEL_L data are taken.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 9

Required additional input:

- RW_FILE=_ <Reactionwheel filename>
- SAMPLE_FREQ=_ <Nominal sample frequency in Hertz>
- The variable LEVEL_H_SOURCE has to be set. If not stated LEVEL_C will be taken as input.

Optional additional input:

- CORR_FACTOR_RW= __ <Factor>. Substituted amplitude is background average multiplied by this factor. Default = 1.1
- DELTA_F_RW= _ <Bandwidth in [Hz]>. Substituted spectral line is $f_{local}\pm$ DELTA_F_RW. Default=0.06 Hz
- FREQ_LOWER_LIMIT= $\sqcup <$ Lowest purged frequency in [Hz]>. Default= 0.15 Hz
- NOT_CORR_TIME_INTERVAL= <Start & End time of interval which is not purged >. Format= hh:mm:ss hh:mm:ss. Several intervals can be stated.

Additionally the disturbances generated by the LAP instrument can be eliminated with this routine.

Optional additional input:

- ELIMINATE_LAP=_ <Y,N> Controls the activation of the purging procedure. Default = Y
- CORR_FACTOR_LAP=_

 Substituted amplitude is background average multiplied by this factor. Default = 1.55
- DELTA_F_LAP= __ <Bandwidth in [Hz]>. Substituted spectral line is $f_{\rm local}\pm {\rm DELTA}_F_LAP.$ Default=0.05 Hz
- LAP_FREQ= $_{\sqcup}$ <Disturbed Frequencies in [Hz] to be eliminated>. Default= 3.2 3.55 6.4 7.1

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 10

NOT_CORR_TIME_INTERVAL= <
 Start & End time of interval which is not purged >.
 Format= hh:mm:ss hh:mm:ss.
 Several intervals can be stated.

Output of the routine are *.LBL and *.TAB files.

3.13 GEN_LEVEL_I_DATA – Generation of averaged and merged LEVEL_H data – (CLI)

This routine is used to merge data of all modes collected on one day, calculate averaged data from this input files and save it as PDS data. As input the RW–corrected LEVEL_H data are taken. The average period has to be stated by the means of the variable RESAMPLED_AVERAGE= $_$. Output of the routine are *.*LBL* and *.*TAB* files.

3.14 GEN_LEVEL_J_DATA – Generation of PCA processed data – (CLJ)

This routine is used to perform a principal component analysis (PCA) and to generate correlated and uncorrelated time series as PDS compliant LEVEL_J data. As input files OB and IB files have to be present in parallel, as the PCA correlates them. Possible input data are averaged LEVEL_G, and LEVEL_I data. The data source has to be specified using the PCA_SOURCE= \Box keyword. Additionally the type of used averaged data has to be stated using the filter keyword FILE_MODE= \Box . E.g FILE_MODE= \Box A16 for processing 16 s averaged data.

Output of the routine are *.LBL and *.TAB files. The output file are split in a CORRELATED and an UNCORRELATED folder. The files in the CORRELATED folder have a suffix $_C$, files in the UNCORRELATED folder end with $_U$.

3.15 GENERATE_PLOTS – Generation of plots

This procedure can be found in the file *PLOT_DAY.PRO* and is used for the automatic plot generation in EPS format. Two general types of plots can be generated:

- Time series plots
- Spectra plots Power spectra and dynamic spectra

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 11

Additionally to the plot generation a $\mathbb{L}T_{E}X$ - shell for every plot is generated by the routine *write_tex*. This routine creates a "figure" environment for every plot (appended to the file d:/ROSETTA/data/plotfiles.tex), so that an easy access to the plot files in a $\mathbb{L}T_{E}X$ - document is provided.

3.16 WRITE_PDS2FTP - The transfer routine of processed PDS data

This routine transfers all recently processed data files (PDS) to the IGEP ftp–server. On the ftp–server a cron job is running, which transfers all new data files once a day to the RPC data server at the Imperial college in London.

3.17 SET_QUALITY – The quality flag application routine

This routine has to be called in the second pass of the data generation process. Invocation is controlled via the keyword SET_QUAL_FLAGS= Y . If this routine is activated, the dummy quality flags (xxxxxxx) of all CALIBRATED, RESAMPLED, and DE-RIVED data which have been written to the files in the first pass of the data generation process will be substituted by the right ones. The right quality information is read from the QUALITY INDEX FILE which has to be specified by the QUAL_FILE= keyword. This ASCII table (*.TXT) is derived from the original EXCEL quality file (*.XLS) by saving the latter one as a MS-DOS formatted text file. The standard index file is =D:\ROSETTA\DATA\QUALITY\RPCMAG_QUALITY_<missionphase>.xls. The index file contains the quality assessment for each sensor and each calibration level for all necessary times. From here the right quality flags are copied to the(*.TAB) data files.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 12

4 Output of DDS2PDS

4.1 Data product overview

DDS2PDS can generate various types of output:

• DATA:

Data output is PDS/PSA compliant. That means for every written data file (ASCII tables, *.TAB) a PDS Label file (*.LBL) will be generated as well. The lines of the label files are automatically formatted to 80 chars (done by $DO_FORMAT_LBL_FILES.PRO$ to be fully PDS/PSA compliant. Additionally to the usual PDS labels geometric labels related to the s/c position at the START_TIME of each file are generated. This is done with the procedure $ADD_GEOINFO_TO_LBL.PRO$

Remark: QA denotes Quality Flags. During the data generation dummy values are written into these columns (done by the procedure $ADD_QUALITY_COLUMN_TO_LBL_TAB.PRO$. The real quality flags have to be set in a second path after the data have been assessed.

The data products are distinguished by processing levels:

- EDITED RAW DATA:

Data in ADC Counts, Magnetic field in instrument coordinates.

FILE	CONTENT				
HK	UTC	OBT	T_OB	T_IB	
	STAGE_ID_A	STAGE_ID_B	FILTER_CFG		
	MAG_REF_VOLT	MAG_NEG_VOLT	MAG_POS_VOLT		
	BX_OB	BY_OB	BZ_OB		
IB & OB	UTC	OBT	BX	BY	ΒZ
	Т	QUALITY			

$ \mathbf{D} \cap \mathbf{C} \mathbf{F} \mathbf{T} \mathbf{T} \mathbf{\Lambda} $ Document: RO-IGEP-	-TR-0007
$ \Pi \cup \Im \Box \Box \Box \Box \Box A $ Issue:	3
Revision:	1
ICED Institut für Geophysik u. extraterr. Physik Date: January	y 20, 2010
IGLI Technische Universität Braunschweig Page:	13

- CALIBRATED DATA:

DATA in physical units, bad vectors removed

* LEVEL_A Data:

Magnetic field in instrument coordinates.

FILE	CONTENT				
HK	UTC	OBT	T_OB	T_IB	
	STAGE_ID_A	STAGE_ID_B	FILTER_CFG		
	MAG_REF_VOLT	MAG_NEG_VOLT	MAG_POS_VOLT		
	BX_OB	BY_OB	BZ_OB		
IB & OB	UTC	OBT			
	BX	BY	BZ	Т	QA

$\ast\,$ LEVEL_B Data:

Magnetic field in instrument s/c – coordinates.

FILE	CONTENT				
IB & OB	UTC OBT				
	BX	BY	BZ	Т	$\mathbf{Q}\mathbf{A}$

* LEVEL_C Data:

Magnetic field in instrument celestial coordinates.

FILE	CONTENT					
IB & OB	UTC OBT POS_X POS_Y POS_Z					
	BX	BY	BZ	QA		

- RESAMPLED DATA:

* LEVEL_K Data:

Magnetic field in instrument $\rm s/c$ – coordinates. Influence of Lander Heater current is corrected

FILE	CONTENT				
IB & OB	UTC OBT				
	BX	BY	BZ	Т	QA

ΡΟς ΕΤΤΛ	Document:	RO–IGEP–TR–0007
$n \cup j \vdash 1 \perp I$	Issue:	3
	Revision:	1
ICDD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLP Technische Universität Braunschweig	Page:	14

* LEVEL_L Data:

Magnetic field in instrument celestial coordinates. Influence of Lander Heater current is corrected

FILE	CONTENT						
IB & OB	UTC	UTC OBT POS_X POS_Y POS_Z					
	BX	BY	BZ	QA			

* LEVEL_E Data:

DATA in physical units, bad vectors removed, data ground calibrated, resampled and averaged to the specified time interval. Source is LEVEL_A data.

Magnetic field in instrument coordinates.

FILE	CONTENT				
IB & OB	UTC	OBT	BX	BY	BZ
	Т	QA			

* LEVEL_F Data:

DATA in physical units, bad vectors removed, data ground calibrated, resampled and averaged to the specified time interval. Source is LEVEL_B or LEVEL_K data.

Magnetic field in s/c-coordinates.

FILE	CONTENT				
IB & OB	UTC	OBT	BX	BY	ΒZ
	Т	$\mathbf{Q}\mathbf{A}$			

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
Technische Universität Braunschweig	Page:	15

* LEVEL_G Data:

DATA in physical units, bad vectors removed, data ground calibrated, resampled and averaged to the specified time interval. Source is LEVEL_C or LEVEL_L data.

Magnetic field in celestial coordinates.

FILE			CONTE	NT	
IB & OB	UTC	OBT	POS_X	POS_Y	POS_Z
	BX	BY	BZ	QA	

* LEVEL_H Data:

Reaction wheel and LAP disturbance corrected Magnetic field in celestial coordinates.

Source is LEVEL_C or LEVEL_L data.

FILE			CONTE	NT	
IB & OB	UTC	OBT	POS_X	POS_Y	POS_Z
	BX	BY	BZ	QA	

* LEVEL_I Data:

Averaged, reaction wheel corrected Magnetic field in celestial coordinates.

FILE	CONTENT				
IB & OB	UTC	OBT	POS_X	POS_Y	POS_Z
	BX	BY	BZ	QA	

- DERIVED DATA:

* LEVEL_J Data:

PCA processed data in celestial coordinates.

Output is one file for the correlated part of the data and two files (IB, OB) containing the uncorrelated "noise" of the original data.

FILE	CONTENT				
_C & _U (IB,OB)	UTC	OBT	POS_X	POS_Y	POS_Z
	BX	BY	BZ	QA	

ROSETTA	Document: Issue	RO–IGEP–TR–0007
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 16

• OFFSET FILES:

The S/W actualizes the two overall offset files

D:\ROSETTA\rpcPds\rpcMag\CALIB\CLA_OFFSET_OB.TXT_

 $\texttt{D:\ROSETTA\rpcPds\rpcMag\CALIB\CLA_OFFSET_IB.TXT}_{\sqcup}$

These files contain entries for the first and the last vector of a *.TAB file. Format:

Time Bx_GND By_GND Bz_GND Bx_FL By_FL Bz_FL OFF_X OFF_y OFF_z T

Time is given as UTC string (yyyy-mm-ddThh:mm:ss), fields in nanotesla and the temperature in Kelvin. Index GND means, that the value is obtained from the Ground Calibration ONLY, index FL assigns the field values after application of the proper inflight calibration. Offset assigns the values which have to be added to the data obtained by the pure inflight calibration to ensure that the first vector of a file fits to the last vector of the previous file.

• PLOTS:

For details refer to section 5.6.2.

• LOGGING FILES:

Each run of DDS2PDS is documented in a logging file named /ROSETTA/DATA/LOG.TXT

• AUTOMATIC MAILS:

The entries of the logging files will be automatically transmitted via mail. The sender and receiver parameters can be specified in the last rows of the $do_pds.pro$ file.

• LAT_EX FIGURE FILES:

For an easy documentation in IAT_{EX} an automatic generation of figure environments is provided by the DDS2PDS s/w. At the first run of DDS2PDS the /ROSETTA/DATA/PLOTFILES.TEX will be generated. For each generated plot a figure environment with caption and labels will be created and written to the file. For each further run of the s/w the file is appended with the new figure blocks. Example:

```
\def\fileA{\special{psfile=D:/Rosetta/rpcPds/rpcMag/work/PLOTS/CALIBRATED/2004/MAY
    /SPECTRA/LEVEL_B/IB/RPCMAG040510T0047_CLB_IB_M3_DS1_10000_002.EPS
    vscale = 80 hscale = 80 angle=0 }}
\begin{figure}[h]
    \begin{figure}[h]
    \begin{center}
    \setlength{\unitlength}{1cm}
    \begin{picture}(15.0,21)(0,0)
    \put(-1.00, 1.00){\fileA }
    \put( -1.00, 1.00){\framebox(17.0,21){}}
    \end{picture}
    \caption{File: RPCMAG040510T0047\_CLB\_IB\_M3\_DS1\_10000\_002}
    \label{RPCMAG040510T0047\_CLB_IB_M3_DS1_10000_002}
    \end{center}
```

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 17

4.2 Special handling of EQMTEST data

From time to time data from the EQM have to be processed. The data and plots shall not be generated in the usual flight model directories but in separate EQM directories. To distinguish both instruments the following steps have to be taken into account to process EQMTEST data:

- Rename the standard PLOTS directory to e.g. PLOTS_FM
- For date of the EQMTEST the MISSION_PHASE_ABBREVIATION has to be set to EQMTEST. This must be done at the specified line in the file SET_GLOBAL_VARIALBES.PRO.
- The CAL_FILE_ID in the CMD files has to be set to GND.
- Process CMD file as usual.
- All generated data will be stored in the ...-EQMTEST-... datasets directories.
- All generated plots appear in the newly created PLOTS directory.
- Move PLOTS to EQMTEST\PLOTS
- Rename the plot directory PLOT_FM back to PLOTS

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 18

5 Directory and Filename Definiton

The root directory of the PDS system and the software is /ROSETTA/rpsPds/rpcMag/

5.1 Software Directories

The software package can be found under ../SOFTWARE/PCWIN/. Here are 4 subdirectories

- *BAT* contains all the batch files for data transfer and conversion.
- *FOR* contains the FORTRAN files from ESA's *OASWLIB* Library to calculate orbit and attitude vectors.
- *IDL* contains all IDL routines.
- *Matlab* contains everything needed for the *Raw2ASCII* conversion

5.2 PDS Data Directories

To avoid any confusion the file names and directories of the output file are generated automatically. The nomenclature follows the PDS rules and the results/definitions of the RPC team meeting in February 2004 and the RPC Archiving Guidelines RO-IGEP-TR00016.

The converted data is stored in the ./DATA directory under the actual base directory. The actual base directory name is directly derived from the DATA_SET_ID and varies for each data product with the mission phase and the product level. This data folder contains different subdirectories sorted primarily by data processing level. Thus there are the

- EDITED
- CALIBRATED
- RESAMPLED
- DERIVED

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 19

Each of these directories contains low level directories sorted by YEAR (yyyy) and MONTH (JAN, FEB, ..., DEC). Below this directory layer the structure varies:

- In the *EDITED* directory there are only 3 data directories for Housekeeping *HK* data, Inboard *IB* and outboard *OB* sensor magnetic field data.
- In the *CALIBRATED* directory an additional distinction is made between different levels:
 - LEVEL_A: Calibrated HK, IB, and OB data in instrument coordinates.
 - LEVEL_B: Calibrated IB, and OB data in s/c-coordinates.
 - LEVEL_C: Calibrated IB, and OB data celestial coordinates.
- The *RESAMPLED* directory contains
 - LEVEL_K:

Calibrated $IB,\,{\rm and}~OB$ data in s/c–coordinates. Heater current influence eliminated.

- LEVEL_L:

Calibrated IB, and OB data celestial coordinates. Heater current influence eliminated.

- $LEVEL_E$ IB, and OB data, which are glued and averaged raw (ADC) data for a complete day.
- LEVEL_F

 $IB, \ {\rm and} \ OB$ data, which are glued and averaged s/c–coordinate data for a complete day.

- LEVEL_G

IB, and OB data, which are glued and averaged celestial coordinate data for a complete day.

- LEVEL_H

 $I\!B\!,$ and $O\!B$ data, which are reaction wheel corrected data in celestial coordinates for a complete day.

- LEVEL_I

 $I\!B,$ and $O\!B$ data, which are glued and averaged LEVEL-H data in celestial coordinates for a complete day.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 20

- The *DERIVED* directory contains up to now
 - LEVEL₋J

 $_C$ correlated data, which are PCA processed, averaged averaged LEVEL_G or LEVEL_I data in celestial coordinates for a complete day.

 $_U$ uncorrelated data (demerged from IB, and OB data), which are PCA processed, averaged LEVEL_G or LEVEL_I data in celestial coordinates for a complete day.

5.3 Plot Directories

The ./*PLOTS* directory has exactly the same structure as the ./*DATA* directory. As the plots will not be part of the PSA/PDS archive the are just stored locally directly under the rpcpds/rpmag/ directory. Add–Ons:

- there is an additional HK_B directory under $LEVEL_A$ containing Magnetic field data from the Housekeeping ADC.
- under the *CALIBRATED* directory there is a *LEVEL_D* folder containing plots for the comparison between the IB and the OB sensor. The *DIF* directory contains plots of the differences between the IB and the OB sensor. The *XXYYZZ* directory contains plots where the i-th component of the IB is plotted versus the i-th component of the OB sensor.
- under the *CALIBRATED*, *RESAMPLED* and *DERIVED* directory there is also a *SPECTRA* folder containing spectrum plots. This directory contains the usual substructure of *LEVEL_n* folder with separate *IB* and *OB* directories.

5.4 Temporary Working Directory

The system needs a temporary directory, where the script file, temporary batch files, and temporary data are stored. This directory is located at d:/ROSETTA/rpcPds/RpcMag/TMP

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 21

5.5 Input, Parameter and RAW2ASCII Directories

All auxiliary files like

- the RAW2ASCII input, output and parameter files,
- the control file *PDS.CMD*,
- the Reaction wheel data files $schk7_{-}yyyy-mm-dd.txt$,
- the Instrument status file $piu_hk_yyyy-mm-dd.txt$
- the heater corrected input files in LEVEL_B format
- $\bullet\,$ and the logfiles

are located in the d:/ROSETTA/data directory and its subdirectories.

5.6 Output filenames

5.6.1 DATA files

The filenames obey the PDS 27.3 standard.

Definition for original sampled and reaction wheel corrected data:

- Magnetic field data: RPCMAG{yymmdd}T{hhmm}_{level}_{sensor}_M{mode}.{extension}
- Hosekeeping data: RPCMAG{yymmdd}T{hhmm}_{level}_HK.{extension}

yymmdd	YEAR, Month, Day: in two digits format
hhmm	Hour, Minute: in two digits format
level	RAW, CLA, CLB, CLC, CLH, CLK, CLL, for EDITED, LEVEL_A,
	LEVEL_B, LEVEL_C, LEVEL_H, LEVEL_K, LEVEL_L
mode	$1 \dots 6$ for SID1 \dots SID6
sensor	OB, IB
extension	LBL for LABEL files, TAB for TABLE files

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 22

Definition for averaged data:

• Magnetic field data: RPCMAG{yymmdd}_{level}_{sensor}_A{avg}.{extension}

yymmdd	YEAR, Month, Day: in two digits format
level	CLE,CLF,CLG,CLI for LEVEL_E, LEVEL_F, LEVEL_G,
	LEVELI
avg	average period in seconds
sensor	OB, IB
extension	LBL for LABEL files, TAB for TABLE files

Definition for PCA processed data:

• CORRELATED Magnetic field data: RPCMAG{yymmdd}_{level}_A{avg}_C.{extension}

yymmdd	YEAR, Month, Day: in two digits format
level	CLJ for LEVEL_J
avg	average period in seconds
extension	LBL for LABEL files, TAB for TABLE files

• UNCORRELATED Magnetic field data: RPCMAG{yymmdd}_{level}_{sensor}_A{avg}_U.{extension}

yymmdd	YEAR, Month, Day: in two digits format
level	CLJ for LEVEL_J
sensor	OB, IB
avg	average period in seconds
extension	LBL for LABEL files, TAB for TABLE files

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 23

5.6.2 PLOT files

Definition:

The name of the plot files is directly derived from the data file name. Thus the beginning of the file names is identical up to the $\{mode\}$ respective $\{HK\}$ item. Behind these characters the structure differs:

- HK and HK_B plots RPCMAG{yymmdd}T{hhmm}_{level}_HK_P{start}_{end}.EPS
- OB and IB plots RPCMAG{yymmdd}T{hhmm}_{level}_{sensor}_M{mode}_T{start}_{end}_{cal_id}.EPS RPCMAG{yymmdd}T{hhmm}_{level}_{sensor}_A{avg}_T{start}_{end}_{cal_id}.EPS
- DIFF plots RPCMAG{yymmdd}T{hhmm}_{level}_M{mode}_DIF_P{start}_{end}_{cal_id}.EPS
- XXYYZZ plots RPCMAG{yymmdd}T{hhmm}_{level}_M{mode}_XXYYZZ_P{start}_{end}_{cal_id}.EPS
- SPECTRA plots RPCMAG{yymmdd}T{hhmm}_{level}_{sensor}_M{mode}_DS{freq_start}_{freq_end}_{cal_id}.EPS RPCMAG{yymmdd}T{hhmm}_{level}_{sensor}_M{mode}_PS{freq_start}_{freq_end}_{cal_id}.EPS RPCMAG{yymmdd}T{hhmm}_{level}_{sensor}_A{avg}_DS{freq_start}_{freq_end}_{cal_id}.EPS RPCMAG{yymmdd}T{hhmm}_{level}_{sensor}_A{avg}_PS{freq_start}_{freq_end}_{cal_id}.EPS
- PCA processed plots RPCMAG{yymmdd}_{level}_A{avg}_C_T{start}_{end}_{cal_id}.EPS RPCMAG{yymmdd}_{level}_{sensor}_A{avg}_U_T{start}_{end}_{cal_id}.EPS
- PCA processed SPECTRA plots RPCMAG{yymmdd}_{level}_A{avg}_C_DS{freq_start}_{freq_end}_{cal_id}.EPS RPCMAG{yymmdd}_{level}_A{avg}_C_PS{freq_start}_{freq_end}_{cal_id}.EPS RPCMAG{yymmdd}_{level}_{sensor}_A{avg}_U_DS{freq_start}_{freq_end}_{cal_id}.EPS RPCMAG{yymmdd}_{level}_{sensor}_A{avg}_U_PS{freq_start}_{freq_end}_{cal_id}.EPS

ROSETTA	Document: Issue:	RO-IGEP-TR-0007 3
IGEP Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
Technische Universität Braunschweig	Page:	24

Nomenclature:

yymmdd	YEAR, Month, Day: in two digits format
hhmm	Hour, Minute: in two digits format
level	CLA,CLB, CLC,CLE, CLF, CLG,CLH,CLI,CLJ for LEVEL_A, LEVEL_B,
	LEVEL_C, LEVEL_E, LEVEL_F, LEVEL_G, LEVEL_H, LEVEL_I, LEVEL_J
mode	$1 \dots 6$ for SID1 \dots SID6
avg	average period in seconds
sensor	OB, IB
Р	means PLOT interval
Т	means TIME series plot
DIF	means DIFFERENCE plot
XXYYZZ	Means Plot of OB and IB components $(x_i^{OB} \text{ vs. } x_i^{IB})$
DS	means DYNAMIC spectrum
\mathbf{PS}	means POWER spectrum
start	Start time of the plot in hhmm–format
end	End time of the plot in hhmm-format
$freq_start$	Start frequency of the plot in mHz
freq_end	End frequency of the plot mHz
cal_id	Version of the calibration files
С	Correlated part of CLJ files
U	Uncorrelated part of CLJ files

5.6.3 Logging files

A logging file of a session is written to D:/ROSETTA/DATA/LOG.TXT. It contains information about the generated files and error messages.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 25

6 The command file *PDS.CMD* and the control parameters of *DDS2PDS*

The DDS2PDS S/W can be completely controlled by the means of the command file PDS.CMD which has to be located at d:/ROSETTA/data. The command file consists of a list of commands which are arranged in command blocks. The file can contain any number of command blocks which are used for a specific tasks, e.g. processing the data of a certain day or just plot data of a specific day. Each block begins with a START: command and is terminated by an END: command. A block is used to process the data a of single day defined by the keyword DAY= \Box The order of keywords within a block is arbitrarily, except for the ACTION: keyword as this will trigger the execution of the data processing.

The file can contain comment lines, starting with a **#**, which will be ignored. Additionally every line will be ignored which does not begin with a defined command keyword. IF a command keyword is stated, it HAS TO HAVE A VALID ARGUMENT!!

6.1 Syntax description

Here is the complete command list including a command description:

• #_u Optional entry!

This character assigns a comment line.

• ACTION:

Necessary Keyword! When the parser detects this keyword, the data processing will be started.

- BOOM=_ Optional Keyword! Valid values:
 - STOWED
 - DEPLOYED
 - any text...

This keyword tells the system the status of the boom.

For data before 2004-03-19T03:35 it HAS to be $\tt STOWED$, after 2004-03-17T03:50 it has to be <code>DEPLOYED</code>.

For the boom deployment phase a text like "boom deployment ongoing" can be used. If the keyword is not used the default value (DEPLOYED) will be used.

ΡΟς ΕΤΤΛ	Document:	RO-IGEP-TR-0007
$\mathbf{N} \cup \mathbf{S} \to \mathbf{I} + \mathbf{I} + \mathbf{A}$	Issue:	3
	Revision:	1
ICDD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
ΙΩΓ Technische Universität Braunschweig	Page:	26

• CAL_DATA=⊔ Necessary Keyword! Valid values:

- N (for No)

- Y (for Yes)

This keyword controls the calibration and PDS file generation process. If enabled, the sorted *.ASS files will be read, ground calibration results will be applied and PDS files will be generated.

• CAL_FILE_ID=_

Necessary Keyword! Valid values:

- GND (for the usage of the ground calibration results only & for generation of a new TEMP_MODEL !)
- <anything>_nnn (STRING + three digit number 001,002,...)
- yyyy-mm-dd_nnn
- yyyy-mm_nnn

The ground calibration files (designator: GND) will automatically be used for generation of the offset–temperature models. For normal use the actual inflight calibration files

 $\label{eq:reconstruction} RPCMAG_<CAL_FILE_ID>_CALIB_OB.TXT \mbox{ and } RPCMAG_<CAL_FILE_ID>_CALIB_IB.TXT$

will be used. They describe a 4th order polynomial temperature model for temperature dependent offset reduction of the magnetic field data. The nnn number (MANDATORY!) of the CAL_FILE_ID= \Box will be used as part of the output plot name for an easy distinction of the plot files.

• CORR_FACTOR_LAP=

Optional Keyword! Valid values:

– Real numbers ≥ 0 , default is 1.1

When frequency lines generated by LAP are eliminated, the new spectral amplitude at the location of the purged line has to be substituted by a definite value. This value is defined by the weighted value of the average amplitude in the time/frequency neighborhood of the actual frequency line. Thus, the new value then will be $average \cdot CORR_FACTOR_LAP$.

Setting the factor to zero will generate a zero amplitude (black line) which is a perfect mean to show the location of the line during the testing phase.

росгт		Document:	RO-IGEP-TR-0007
NUSEI	$\perp A$	Issue:	3
		Revision:	1
ICDD Institut f	ür Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLI Techn	ische Universität Braunschweig	Page:	27

● CORR_FACTOR_RW=_

Optional Keyword! Valid values:

- Real numbers ≥ 0 , default is 1.55

When frequency lines generated by the Reaction wheels are eliminated, the new spectral amplitude at the location of the purged line has to be substituted by a definite value. This value is defined by the weighted value of the average amplitude in the time/frequency neighborhood of the actual frequency line. Thus, the new value then will be

 $average \cdot CORR \ FACTOR \ RW.$

Setting the factor to zero will generate a zero amplitude (black line) which is a perfect mean to show the location of the line during the testing phase.

• DAY=

Necessary Keyword! Format: yyyy-mm-dd This keyword specifies the day to be analyzed.

● DEL=⊔

Optional Keyword! Format: filename to be deleted. Multiple wildcards allowed.

Using this keyword (multiple entries possible) files can be deleted. This is useful, if in the automatic generation process some files were produced which are not of interest. E.g. after the processing of the data from the 7th of May 2004 in all modes it might be possible that there are only a few vectors in the mode M1 IB files which shall not be archived. They will be deleted using $DEL=_{\sqcup}*040507*IB_M1*$. The routine deletes only files in the working directory maindir =D:/rosetta/rpcdps/rpcmag and its subdirectories. Preferably the $DEL=_{\sqcup}$ command should be located behind the ACTION: command, so that the just generated files can be deleted and do not consume any space on the disk anymore.

• DELTA_F_LAP=⊔ Optional Keyword! Valid values:

- Real numbers ≥ 0 , default is 0.06

This keyword defines the bandwidth (frequency in Hz) of the LAP spectral line to be eliminated. All amplitudes within an interval of $\pm DELTA_F_LAP$ around the actual local frequency will be purged.

	Document:	RO–IGEP–TR–0007
$\Pi \cup \Im \Box \Box \Box \Box A$	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
ТССГ Technische Universität Braunschweig	Page:	28

- DELTA_F_RW=⊔ Optional Keyword! Valid values:
 - Real numbers ≥ 0 , default is 0.05

This keyword defines the bandwidth (frequency in Hz) of the Reaction Wheel spectral line to be eliminated. All amplitudes within an interval of $\pm DELTA_F_RW$ around the actual local frequency will be purged.

 DYNSPEC_WIN_WIDTH=□ Mandatory, if GEN_DYNSPEC_PLOTS=□Y Valid values:

- Integer number $1 \dots 16$, default is 7.

The window width is used for the calculation of dynamic spectra. The higher this value, the sharpener is the frequency resolution. However, with increasing windows the time resolution decreases.

The default window width is 7, hence $2^7 = 128$.

 $11~\mathrm{is}$ a good value for SID3 and SID4 data.

8 (short time interval) or 11 (long time interval) is good for SID2 data.

• ELIMINATE_LAP=⊔ Optional Keyword!

Valid values:

- N (for No)

- Y (for Yes, DEFAULT if GEN_LEVEL_H_DATA=_Y)

This keyword controls the data cleaning with respect to the disturbance caused by the LAP instrument. Correction can be enabled (DEFAULT) for the LEVEL_H data generation. If enabled the constant frequency lines generated by LAP will be purged.

● ELIMINATE_RW=_

Optional Keyword! Valid values:

- N (for No)

- Y (for Yes, DEFAULT if GEN_LEVEL_H_DATA= $_{\sqcup}$ Y)

This keyword controls the data cleaning with respect to the disturbance caused by the ROSETTA Reaction wheels. Correction can be enabled (DEFAULT) for the LEVEL_H data generation. If enabled the varying frequency lines generated by the RW will be purged.

BOSETTA	Document:	RO-IGEP-TR-0007
	Issue:	3
	Revision:	1
ICTD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLI Technische Universität Braunschweig	Page:	29

• END:

Necessary Keyword! This keyword terminates a block and separates it from other ones.

• FILE= Optional Keyword! Valid values:

- Complete path of a low level command file, e.g. D:/Rosetta/data/test.cmd

This keyword can be used to maintain a high level structure in the command file. If this keyword is used, the command parser will execute the commands in the stated sub command file (here: test.cmd). No other keywords should be used in the high level *PDS.CMD* file if FILE= \Box is used. As the usage of FILE= \Box will lead to a recursive process, the FILE= \Box keyword should not be used in the low level command files!

● FILE_MODE=_

Optional Keyword! Valid values:

- M1 - M2 - M3 - M4
- M5
- -M6
- OB
- IB

For a higher state of filtering the value can be extended to a combined sensor_mode filter or an average selector like e.g

- OB_M1
- IB_M3
- A16
- ...

This keyword acts as a filter. If specified, only data of the stated mode will be processed (Input File Path Filter).

If keyword is omitted, all data will be processed. Especially for generating PCA processed LEVEL_J data the average interval like FILE_MODE= $\Box A1$ for 1s avg data has to be stated.

R O S E T T A	Document:	RO–IGEP–TR–0007
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 30

• FILE_SELECT=

Optional Keyword! Valid values:

- D:\ROSETTA\rpcPDS\rpcMag\work\data\....
- D:\ROSETTA\rpcPDS\rpcMag\work\data\RPCMAG*050302*CLA*OB*

- ...

This keyword acts as a filter. If specified, only data files matching the stated expression will be processed.

This keyword acts on the highest filter level and will override the FILE_MODE= \Box in the first filter step and the automatic selection if FILE_SELECT= \Box contains more then 4 characters. In this case FILE_MODE= \Box is NOT taken into account. So use this keyword with care! If used, the complete path down to root level should be named. If used, also the keywords INSTRUMENT= \Box and FILE_MODE= \Box should be set to guarantee a successful filtering.

● FILE_TIME_END=_

Optional Keyword!

Format: hh:mm:ss

If this keyword is present, only data before the given time will be processed. Otherwise (default) all data are taken into account.

● FILE_TIME_START=_

Optional Keyword!

Format: hh:mm:ss

If this keyword is present, only data after the given time will be processed. Otherwise (default) all data are taken into account.

● FREQ_LOWER_LIMIT=_

Optional Keyword!

Valid values:

Format: Real numbers ≥ 0 , default is 0.15

This keyword defines the lowest frequency (frequency in Hz) which will be purged during the elimination of the Reaction wheel disturbances. Below the limit frequency the the reaction wheel disturbance can usually not be identified as the spectral density of the real signal is much higher. A fine tuning of this value might be favorably for SID4 data.

ΡΟς ΓΤΤΛ	Document:	RO–IGEP–TR–0007
	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGEF Technische Universität Braunschweig	Page:	31

- GEN_DYNSPEC_PLOTS=∟ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

Using this keyword the generation of dynamic spectra plots is enabled. In this case further keywords as SPEC_SOURCE= $_{\sqcup}$, PLOT_FREQ_START= $_{\sqcup}$, PLOT_FREQ_END= $_{\sqcup}$, PLOT_AMP_MIN= $_{\sqcup}$, and PLOT_AMP_MAX= $_{\sqcup}$ have to be set (see below). With these keywords detailed plot parameters can be adjusted.

- GEN_LEV_A_PLOTS=_ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

If enabled HK, IB and OB plots of the CALIBRATED LEVEL_A data will be generated.

- GEN_LEV_B_PLOTS=_ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

If enabled IB and OB plots of the CALIBRATED LEVEL_B data will be generated.

- GEN_LEV_C_PLOTS= Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

If enabled IB and OB plots of the CALIBRATED LEVEL_C data will be generated.

ΡΟς ΕΤΤΛ	Document:	RO–IGEP–TR–0007
$n \cup s \perp 1 \perp A$	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
ТGLГ Technische Universität Braunschweig	Page:	32

• GEN_LEV_D_PLOTS=⊔ Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

If enabled difference plots of IB and OB (DIFF) and layer plots (XXYYZZ) plots of the CALIBRATED LEVEL_B data will be generated.

• GEN_LEV_E_DATA=_ Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

This keyword allows the generation of averaged data. As input the EDITED RAW PDS data has to be available. All files (all different modes) of the stated day will be glued together and averaged PDS data will be generated. The average period has to be assigned by the RESAMPLED_AVERAGE= $_{\sqcup}$ keyword.

• GEN_LEV_E_PLOTS=⊔ Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

If enabled a plot from the merged and averaged LEVEL_E data will be generated. This is useful especially for the commissioning phase where lots of mode changes happened during one session.

• GEN_LEV_F_DATA=_ Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

This keyword allows the generation of averaged data. As input the LEVEL_B or LEVEL_K PDS data has to be available. All files (all different modes) of the stated day will be glued together and averaged PDS data will be generated. The average period has to be assigned by the RESAMPLED_AVERAGE= $_{\sqcup}$ keyword. The data source is selected via the LEVEL_F_SOURCE= $_{\sqcup}$ keyword.

ΒΟςΕΤΤΛ	Document:	RO–IGEP–TR–0007
	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLI Technische Universität Braunschweig	Page:	33

- GEN_LEV_F_PLOTS=⊔ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

If enabled a plot from the merged and averaged LEVEL_F data will be generated. This is useful especially for the commissioning phase where lots of mode changes happened during one session.

- GEN_LEV_G_DATA=_ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

This keyword allows the generation of averaged data. As input the LEVEL_C or LEVEL_L PDS data has to be available. All files (all different modes) of the stated day will be glued together and averaged PDS data will be generated. The average period has to be assigned by the RESAMPLED_AVERAGE= $_{\sqcup}$ keyword. The data source is selected via the LEVEL_G_SOURCE= $_{\sqcup}$ keyword.

• GEN_LEV_G_PLOTS=⊔ Optional Keyword! Valid values:

N (for No)
Y (for Yes)

If enabled a plot from the merged and averaged LEVEL_G data will be generated. This is useful especially for the commissioning phase where lots of mode changes happened during one session.

- GEN_LEV_H_DATA=_ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

This keyword allows the generation of reaction wheel corrected data. As input the LEVEL_C or LEVEL_L PDS data has to be available. Only one mode and sensor can be handled within one START/END block. Additional the nominal sample frequency SAMPLE_FREQ= $_{\sqcup}$ the RW file name RW_FILE= $_{\sqcup}$ and the extended file mode (e.g. 'OB_M3' or 'IB_M2' etc.) keyword FILE_MODE= $_{\sqcup}$ has to be defined. The data source is selected via the LEVEL_H_SOURCE= $_{\sqcup}$ keyword.

	Document:	RO–IGEP–TR–0007
$\Pi \cup \Im \Box \square \square H$	Issue:	3
	Revision:	1
ICDD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
ТGLГ Technische Universität Braunschweig	Page:	34

• GEN_LEV_H_PLOTS=⊔ Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

If enabled a plot from the RW corrected LEVEL_H data will be generated.

• GEN_LEV_I_DATA= Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

This keyword allows the generation of averaged data. As input the LEVEL_H PDS data has to be available. All files (all different modes) of the stated day will be glued together and averaged PDS data will be generated. The average period has to be assigned by the RESAMPLED_AVERAGE= \Box keyword.

• GEN_LEV_I_PLOTS=⊔ Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

If enabled a plot from the merged and averaged LEVEL_I data will be generated.

• GEN_LEV_J_DATA=⊔ Optional Keyword!

Valid values:

- N (for No)
- Y (for Yes)

This keyword allows the generation of PCA processed data. As input the LEVEL_G or LEVEL_I PDS data has to be available. PDS data will be generated. The average period has to be assigned by the FILE_MODE=An_{\cup} keyword where n denotes the seconds of the average interval. It is mandatory to state the source of the averaged data with the keyword PCA_SOURCE=LEVEL_G_{\cup} or PCA_SOURCE=LEVEL_I_{\cup}

ΡΟς ΓΤΤΑ	Document:	RO-IGEP-TR-0007
$n \cup j \vdash 1 \perp A$	Issue:	3
	Revision:	1
ICDD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGEP Technische Universität Braunschweig	Page:	35

- GEN_LEV_J_PLOTS=⊔ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

If enabled a plots from the PCA processed LEVEL_J data will be generated.

- GEN_LEV_K_DATA=_ Optional Keyword! Valid values:
 - N (for No, DEFAULT)
 - Y (for Yes)

This keyword allows the generation heater influence corrected data in s/c-coordinates. As input the LEVEL_B (*.TAB) data in the D:\ROSETTA\DATA\HEATER directory has to be available.

- GEN_LEV_K_PLOTS= Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

If enabled a plot from the heater current corrected LEVEL_K data will be generated.

- GEN_LEV_L_DATA=_ Optional Keyword! Valid values:
 - N (for No, DEFAULT)
 - Y (for Yes)

This keyword allows the generation heater influence corrected data in ECLIPJ2000 coordinates. As input the LEVEL_K data has to be available.

- GEN_LEV_L_PLOTS=⊔ Optional Keyword! Valid values:
 - N (for No)
 - Y (for Yes)

If enabled a plot from the heater current corrected LEVEL_L data will be generated.

ΡΟς ΓΤΤΑ	Document:	RO–IGEP–TR–0007
$n \cup s \vdash 1 \perp A$	Issue:	3
	Revision:	1
IODD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLP Technische Universität Braunschweig	Page:	36

● GEN_PLOTS=_

Optional Keyword! Valid values:

- N (for No)
- Y (for Yes)

This keyword controls the automatic plot generation. If enabled the overall plot routine access is activated and EPS files can be produced. The kinds of plots to be produced have to be specified by further keywords (see below). If disabled no plots are generated.

• GEN_POW_SPEC_PLOTS=⊔ Optional Keyword! Valid values:

- N (for No)

- Y (for Yes)

Using this keyword the generation of Power spectra plots is enabled. In this case further keywords as SPEC_SOURCE= $_{\sqcup}$, PLOT_FREQ_START= $_{\sqcup}$, PLOT_FREQ_END= $_{\sqcup}$, PLOT_AMP_MIN= $_{\sqcup}$, and PLOT_AMP_MAX= $_{\sqcup}$ have to be set (see below). With these keywords detailed plot parameters can be adjusted.

• GEN_RAW_PLOTS=⊔ Optional Keyword! Valid values:

> - N (for No) - Y (for Yes)

If enabled HK, IB and OB plots of the EDITED data will be generated.

$D \cap C F T T A$	Document:	RO–IGEP–TR–0007
$n \cup s \vdash 1 + A$	Issue:	3
	Revision:	1
ICDD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGEL Technische Universität Braunschweig	Page:	37

• GEN_TEMP_MODEL=⊔ Optional keyword!

Valid values:

- N (for No)
- Y (for Yes)

If enabled the files needed for the evaluation of the offset–temperature model will be created. This means that the data will be generated using the ground calibration result files

RPCMAG_GND_CALIB_FSDPU_FMOB.TXT

RPCMAG_GND_CALIB_FSDPU_FMIB.TXT

By the means of the keyword $CAL_DATA=_{\sqcup}$ it has to be stated whether the needed LEVEL_A files in instrument coordinates shall be produced. This should be done if not already done. As a last step the LEVEL_E_DATA will be generated from the LEVEL_A_DATA using the average period stated by the RESAMPLED_AVERAGE=_{\sqcup} keyword. 200 s should be used for this parameter. At the end the LEVEL_A and LEVEL_E data are available in the $D:/ROSETTA/DATA/CALIB/TEMPMODEL_AVG$ directory, where they are expected by the modelling s/w $CALIB_ROS_TEMP_XXXX.PRO$ (refer to section10.1) In case of generating a TEMP_MODEL the CAL_FILE_ID has to be set to GND!

● GET_DATA=⊔

Necessary Keyword! Valid values:

- N (for No)
- Y (for Yes)

This keyword controls the data transfer from the IGEP ftp–server to the local machine. If enabled the data will be ftped to the local machine and RAW2ASCII will generate ASCII raw files. If disabled the sorted ASCII raw files have already to be available in the $D:/ROSETTA/DATA/ASS_FILES$ directory.

- INSTRUMENT=⊔
 Optional Keyword!
 Valid values:
 - RPCMAG
 - ROMAP

This keyword specifies the instrument whose is data will be analyzed. If not stated, RPCMAG will be used as default.

ΡΟΥΓΤΑ	Document:	RO-IGEP-TR-0007
$\Pi \cup \Im \Box \square \square \square A$	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLI Technische Universität Braunschweig	Page:	38

- LAP_FREQ=□ Optional Keyword! Valid values: Format: real numbers [Hz] separated by □ Default: 3.2 3.55 6.4 7.1 The stated frequencies denote the disturbed spectral lines to be purged during the LAP correction.
- LEVEL_F_SOURCE=∟ Mandatory, if GEN_LEV_F_DATA=⊔Y Valid values:
 - LEVEL_B, Default
 - LEVEL_K

This keyword specifies which data are used for the generation of LEVEL_F data.

- LEVEL_G_SOURCE=∟ Mandatory, if GEN_LEV_G_DATA=∟Y Valid values:
 - LEVEL_C, Default
 - LEVEL_L

This keyword specifies which data are used for the generation of LEVEL_G data.

- LEVEL_H_SOURCE=_ Mandatory, if GEN_LEV_H_DATA=_Y Valid values:
 - LEVEL_C, Default
 - LEVEL_L

This keyword specifies which data are used for the generation of LEVEL_H data.

• NOT_CORR_TIME_INTERVAL=_

Optional Keyword!

Format: start(hh:mm:ss) end(hh:mm:ss)

This keyword only affects the LEVEL_H data generation. With this keyword the time interval (start time and end time) of non operation periods can be set. Thus, within such a defined interval, the data will NOT be cleaned. The keyword can be repeated several times to define more than one interval where no cleaning action is done. Especially at Earth Swingbys, where the field is increasing, the periods of high magnitudes should not be manipulated as this could generate ugly side effects.

ROSETTA	Document: Issue:	RO-IGEP-TR-0007
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 39

- PCA_SOURCE= Mandatory, if GEN_LEV_J_DATA=__Y Valid values:
 - LEVEL_G
 - LEVEL_I

This keyword specifies which data are used for the PCA.

• PLOT_AMP_MAX=

Mandatory, if GEN_DYNSPEC_PLOTS=_Y or GEN_POW_SPEC_PLOTS=_Y

Format : value in $\frac{nT^2}{Hz}$ (e.g. 1e0) This keyword specifies the upper limit of the spectral density axis plot range in the spectra plots.

• PLOT_AMP_MIN=

Mandatory, if GEN_DYNSPEC_PLOTS=_Y or GEN_POW_SPEC_PLOTS=_Y Format : value in $\frac{nT^2}{Hz}$ (e.g. 1e-4) This keyword specifies the lower limit of the spectral density axis plot range in the

spectra plots. • PLOT_FREQ_END= Mandatory, if GEN_DYNSPEC_PLOTS=_Y or GEN_POW_SPEC_PLOTS=_Y Format : value in mHz (e.g 1000) This keyword specifies the upper limit of the frequency axis plot range in the spectra

plots.

- PLOT_FREQ_START= Mandatory, if GEN_DYNSPEC_PLOTS=_Y or GEN_POW_SPEC_PLOTS=_Y Format : value in mHz (e.g. 0) This keyword specifies the lower limit of the frequency axis plot range in the spectra plots.
- PLOT_TIME_END= Mandatory, if GEN_PLOTS=_Y Format: hh:mm:ss This keyword specifies the upper limit of the time axis plot range.
- PLOT_TIME_START= Mandatory, if GEN_PLOTS=_Y Format: hh:mm:ss This keyword specifies the lower limit of the time axis plot range.

ΡΟς Ε Τ Τ Δ	Document:	RO–IGEP–TR–0007
$n \cup j \vdash 1 \perp I$	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
ТССГ Technische Universität Braunschweig	Page:	40

• QUAL_FILE=_ Mandatory, if SET_QUAL_FLAGS=_Y

Valid values:

 Filename of the actual source file for the quality flag data. The complete name including path has to be stated.

• REFERENCE_BODY=

Optional Keyword! Valid values:

- CHURYUMOV-GERASIMENKO
- EARTH
- LUTETIA
- MARS
- STEINS
- SUN

This keyword defines the used coordinate system center. If the keyword is not set, the default, 'SUN', is taken. In the case of 'SUN' the DDS ORHR files will be used to retrieve Orbit information. If 'EARTH' is set, the ORER files will be used, and if 'MARS' is set, then the ORMR files are used. The keyword influences the LEVEL_C (and higher) data and plot output.

• RESAMPLED_AVERAGE=

Mandatory, if <code>GEN_LEV_E_DATA=_Y</code> or <code>GEN_LEV_F_DATA=_Y</code> or <code>GEN_LEV_G_DATA=_Y</code> Valid values:

- number (in seconds)

This keyword defines the period over which the data shall be averaged.

• RW_FILE=

Mandatory, if GEN_LEV_H_DATA=__Y Valid values:

- Filename of the actual Reaction wheel frequency file. The complete name including path has to be stated.
- SAMPLE_FREQ=_
 - Mandatory, if GEN_LEV_H_DATA=_Y Valid values:
 - Frequency values in Hertz.

ΡΟςεττλ	Document:	RO–IGEP–TR–0007
$\mathbf{N} \cup \mathbf{O} \to \mathbf{E} + \mathbf{I} + \mathbf{A}$	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLI Technische Universität Braunschweig	Page:	41

This keyword informs the s/w about the actual NOMINAL sampling frequency. This is needed for the RW-Correction algorithm as there is no automatic possibility to decide which sensor is the primary one. So, e.g. for SID3 OB as primary define $SAMPLE_FREQ=_120.0$, for SID3 IB as secondary sensor use $SAMPLE_FREQ=_11.0$. The s/w takes the difference between nominal and real sampling frequency into account.

- SET_QUAL_FLAGS= Optional Keyword! Valid values:
 - Y: Quality flags will be updated in TAB files
 - N: No action related to quality flags

The file selection of the TAB files to be updated is done via the DAY, FILE_MODE, and FILE_SELECT keywords. The FILE_SELECT has the highest priority and will be used exclusively if the string length of file mode value exceeds 5 characters. The selection is done using the following command

```
\texttt{selection=}_{\sqcup}\texttt{file\_selection}_{\sqcup}
```

If not stated or the length is below 5 chars, the file selection is done via DAY and FILE_MODE in the following way:

```
selection=_MAIN_DIR+'*'+today_year_+today_month+today_day+'*'+file_mode+'*.TAB'_
```

If no FILE_MODE is set, all files of the assigned day will be processed.

● SMOOTHING_WIN=

Optional Keyword! Valid values:

- 0: No smoothing
- $-3\ldots 15$, value has to be odd

This keyword controls the smoothing of the spectra output. A value of 0 means, that the output picture will not be smoothed. The higher the value, the higher the smoothing effect. 11 is the default value. 11 is good for SID3 data. For SID2 data a value of 5 or 0 is reasonable. SID4 is best with 0.

ΡΟς ΕΤΤΛ	Document:	RO–IGEP–TR–0007
$\mathbf{N} \cup \mathbf{O} \to \mathbf{E} + \mathbf{I} + \mathbf{A}$	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLI Technische Universität Braunschweig	Page:	42

● SPEC_SOURCE=_

Mandatory keyword, if <code>GEN_DYNSPEC_PLOTS=_Y</code> or <code>GEN_POW_SPEC_PLOTS=_Y</code> Valid values:

- LEVEL_A
- LEVEL_B
- LEVEL_C
- LEVEL_K
- LEVEL_L
- LEVEL_E
- LEVEL_F
- LEVEL_G
- LEVEL_H
- LEVEL_I
- LEVEL_J

This keywords specifies the input data for the spectrum plots. All calibrated data level can be chosen to act as the source for the desired plots.

• START:

Necessary Keyword!

This keyword has to be the first command of a block and starts a command block definition.

ΡΟς ΓΤΤΛ	Document:	RO–IGEP–TR–0007
$\mathbf{N} \cup \mathbf{S} \to \mathbf{I} + \mathbf{I} + \mathbf{A}$	Issue:	3
	Revision:	1
IODD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLP Technische Universität Braunschweig	Page:	43

6.2 Example of a Command File

# FILE: PDS.CMD	
# Command file for automatic PDS & P	lot generation
# A block is delimited by START: an	d END: .
# The File can contain several block	S
# Description	
#START:	; Beginning of Block
<pre>#REFERENCE_BODY= EARTH</pre>	; SUN, MARS; coord. Sys. center; select. for ORHR, ORER, ORMR files
#INSTRUMENT=	; RPCMAG or ROMAP - if not stated RPCMAG will be used
#CAL_FILE_ID= nnn	; GND or nnn,001,002,calfile_name for desired cal file designator
#GEN_TEMP_MODEL=	; normally N, if Y temperature model files will be created
#DAY= vvvv-mm-dd	: day to be analyzed
#BOOM= STOWED.DEPLOYED.TEXT	; not necessary. DEPLOYED is default
#FILE= D:\Rossetta\data\test.cmd	: low level command file to be executed
#FILE MODE= M3	; only mode Mx data will be generated or plotted not necessary
#FILE MODE= OB M3	; only mode OB M3 data will be generated or plotted not necessary
#FILE SELECT= D:\ROSETTA\rpcPDS\rpCm	ag/work/data/RPCMA040909T0100*CIC*0B*
#FILE TIME START= 10:00:00	start time of data used ontional
#FILE_TIME_END= 22:00:00	; and time of data used, optional
#CET DATA = N / V	, Cot data from fits gerwar and gtart MATIAR
#GAI DATA = N / Y	, Get data from fip server and start mailed
#CRL_DATA- N / I #CEN LEV E DATA- N /V	, read MAILAD Output and generate FDS files
#GEN_LEV_E_DATA- N /I	; average data from ALL modes
#GEN_LEV_F_DATA- N /I #GEN_LEV_C_DATA- N /V	; average data from ALL modes
#GEN_LEV_G_DATA N /I	, average data from ALL modes
#GEN_LEV_H_DATA= N /Y	; generate KW corrected data
#GEN_LEV_I_DATA= N /Y	; average LEVEL_H data of all modes
#GEN_LEV_J_DATA= N /Y	; process PCA
#GEN_LEV_K_DATA= N /Y	; generate heater influence corrected data in s/c coords.
#GEN_LEV_L_DATA= N /Y	; generate heater influence corrected data in ECLIPJ2000 coords.
#RESAMPLED_AVERAGE= <seconds></seconds>	;
#GEN_PLOTS= Y / N	; generate plots
#GEN_RAW_PLOTS= Y / N	; gen. RAW data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_A_PLOTS= Y/N	; gen. LEVEL_A data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_B_PLOTS= Y/N	; gen. LEVEL_B data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_C_PLOTS= Y/N	; gen. LEVEL_C data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_D_PLOTS= Y/N	; gen. LEVEL_D data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_E_PLOTS= Y/N	; gen. LEVEL_E data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_F_PLOTS= Y/N	; gen. LEVEL_F data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_G_PLOTS= Y/N	; gen. LEVEL_G data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_H_PLOTS= Y/N	; gen. LEVEL_H data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_I_PLOTS= Y/N	; gen. LEVEL_I data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_J_PLOTS= Y/N	; gen. LEVEL_J data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_K_PLOTS= Y/N	; gen. LEVEL_K data plots, only necessary if GEN_PLOTS=Y
#GEN_LEV_L_PLOTS= Y/N	; gen. LEVEL_L data plots, only necessary if GEN_PLOTS=Y
#GEN_DYNSPEC_PLOTS= N	; gen. SPECTRUM plots, only necessary if GEN_PLOTS=Y
#LEVEL_F_SOURCE= LEVEL_B /LEVEL_K	; input source for LEVEL_F data
#LEVEL_G_SOURCE= LEVEL_C /LEVEL_L	; input source for LEVEL_G data
#LEVEL_H_SOURCE= LEVEL_B /LEVEL L	; input source for LEVEL_H data
#DYNSPEC_WIN_WIDTH= 7	; window width as 2 ⁿ e.g. 2 ⁷ ; increas. n sharpens the f resolution
#SMOOTHING WIN= 0	: smoothing of spectra, 0.115, odd. The higher the more smoothed
#GEN POW SPEC PLOTS= Y /N	; gen. POWER SPECTRUM plots, only necessary if GEN PLOTS=Y
#SPEC SOURCE= LEVEL B	: LEVEL A.LEVEL B.LEVEL C.LEVEL E.LEVEI. F.LEVEI. G.LEVEI. H.LEVEI. T
	, , , _ , _ , , _ , , _ , , , _ ,

```
ROSETTA
                                                           Document: RO-IGEP-TR-0007
                                                           Issue:
                                                                                           3
IGEP Institut für Geophysik u. extraterr. Physik
                                                           Revision:
                                                                                           1
                                                           Date:
                                                                           January 20, 2010
                Technische Universität Braunschweig
                                                           Page:
                                                                                          44
  #PCA_SOURCE= LEVEL_G
                                     ; LEVEL_G or LEVEL_I
  #PLOT_TIME_START= hh:mm:ss
                                     ; necessary if GEN_PLOTS=Y
  #PLOT_TIME_END= hh:mm:ss
                                     ; necessary if GEN_PLOTS=Y
                                     ; start frequency in mHz
  #PLOT_FREQ_START= 0
  #PLOT_FREQ_END= 0
                                     ; end frequency in mHz
  #PLOT_AMP_MIN= 1e-4
  #PLOT_AMP_MAX= 1e0
  #RW_FILE= D:\Rossetta\data\REACTION_WHEEL\DATA\ASCII\rwheel_2004_SEP0710.txt ; rw file name
  #SAMPLE_FREQ= 20.0
                                     ; sampling frequency in Hertz - used for RW correction
  #ELIMINATE_LAP= Y/N
                                     ; LAP disturbance correction at LEVEL_H data generation, Y=Default
  #ELIMINATE_RW= Y/N
                                    ; RW disturbance correction at LEVEL_H data generation, Y=Default
  #CORR_FACTOR_RW= 1.1
                                    ; average * Factor give new spectral amplitude
  #CORR_FACTOR_LAP= 1.55 ; average * Factor give new spectral amplitude
#LAP_FREQ= 3.2 3.55 6.4 7.1 ; Frequency lines to be purged at LAP correction
  #NOT_CORR_TIME_INTERVAL= 12:00:00 13:00:00 ; Time intervals in which no correction is performed
  #NOT_CORR_TIME_INTERVAL= 15:30:00 16:45:00 ;
  #NOT_CORR_TIME_INTERVAL= 22:30:00 23:15:00 ;
                                     ; Bandwidth of RW correction is 2* DELTA_F_RW [Hz]
  #DELTA_F_RW= 0.06
  #DELTA_F_LAP= 0.05
                                    ; Bandwidth of LAP correction is 2* DELTA_F_LAP [Hz]
  #FREQ_LOWER_LIMIT=0.15
                                     ; Lowest RW frequency that will be corrected
  #QUAL_FILE= D:\ROSETTA\DATA\QUALITY\quality_index.txt
                                     ; quality flags to be updated in TAB file. Source is QUAL_FILE
  #SET_QUAL_FLAGS= Y/N
  #DEL=
                                     ; file to be deleted maindir =D:\rosetta\rpcdps\rpcmag
  #ACTION:
                                      ; Start processing
  #END:
  #
  #
  ****************
  START:
  DAY= 2004-03-17
  CAL_FILE_ID= 002
  BOOM= STOWED
  GET_DATA= N
  CAL_DATA= N
  GEN_LEV_E_DATA= Y
  GEN_PLOTS= Y
  GEN_RAW_PLOTS= N
  GEN_LEV_A_PLOTS= Y
  GEN_LEV_B_PLOTS= N
  GEN_LEV_C_PLOTS= N
  GEN_LEV_D_PLOTS= N
  GEN_LEV_E_PLOTS= Y
  GEN_LEV_K_PLOTS= Y
  GEN_LEV_L_PLOTS= Y
  RESAMPLED_AVERAGE= 60
  GEN_DYNSPEC_PLOTS= N
  GEN_POW_SPEC_PLOTS= N
  SPEC_SOURCE= LEVEL_A
  PLOT_TIME_START= 22:50:00
  PLOT_TIME_END= 24:00:00
  PLOT_FREQ_START =
  PLOT_FREQ_END =
```

ΡΟςεττλ	Document:	RO–IGEP–TR–0007
$\mathbf{N} \cup \mathbf{S} \to \mathbf{I} \to \mathbf{A}$	Issue:	3
	Revision:	1
ICTD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLГ Technische Universität Braunschweig	Page:	45

PLOT_AMP_MIN= 1e-4 PLOT_AMP_MAX= 1e0 ACTION: END:

START: DAY= 2004-03-19 CAL_FILE_ID= 002 BOOM= STOWED FILE_MODE= M3 GET_DATA= N CAL_DATA= N GEN_PLOTS= Y GEN_RAW_PLOTS= N GEN_LEV_A_PLOTS= N GEN_LEV_B_PLOTS= Y GEN_LEV_C_PLOTS= N GEN_LEV_D_PLOTS= Y GEN_DYNSPEC_PLOTS= N GEN_POW_SPEC_PLOTS= Y SPEC_SOURCE= LEVEL_B PLOT_TIME_START= 01:30:00 PLOT_TIME_END= 03:30:00 PLOT_FREQ_START= 1000 PLOT_FREQ_END= 10000 PLOT_AMP_MIN= 1e-4 PLOT_AMP_MAX= 1e0 ACTION: END:

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3			
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 46			

7 File format of *RAW2ASCII* files

7.1 Input files

As input files the original ROSETTA TM files from the DDS are used. They are retrieved automatically via an RSYNC process from the ROSETTA data server XANTHUS at Imperial College and stored on the IGEP-FTP server.

7.2 Output files

RAW2ASCII converts one input file to various output files. These files are split up by sensor and mode.

7.2.1 Housekeeping files

FORMAT										
Filename:	RPO	RPCMAG{yymmdd}T{hhmm}_RAW_HK.ASC								
Header	#	#								
Datacolumn	1.	Timestring (SCET in UTC)	$\{yyyy-mm-dd\}T\{hh\}:\{mm\}:\{ss\}.\{fffff\}$							
	2.	S/C CLOCK OBT	{nnnnnnnnn}.{nnnnn}							
	3.	Temperature OB	ADC Value							
	4.	Temperature IB	ADC Value							
	5.	Stage A ID	Value							
	6.	Stage B ID	Value							
	7.	Filter Config	Value							
	8.	MAG Ref. Voltage	ADC Value							
	9.	MAG Neg. Supply Volt.	ADC Value							
	10.	MAG Pos. Supply Volt.	ADC Value							
	11.	MAG Bx (OB)	ADC Value							
	12.	MAG By (OB)	ADC Value							
	13.	MAG Bz (OB)	ADC Value							

Example:

#	FILE=	RPCMAG	0403171	C2257_R	RAW_HK.AS	С											
#	SCET (OBT TEM	POB TEM	IPIB ST	TAGEA-ID	STAGEB-ID	FILTERCFG	MAGREFVOLTA	GE MAGUB-	MAGUB+	MAGX	MAGY MAGZ					
20	04-03-	-17T22:	58:30.5	36251	38185096	.99753	12441	12766	1	2	0	262117	252	31	65181	65382	2236
20	04-03-	-17T22:	59:02.5	36262	38185128	.99753	12447	12768	1	2	0	262120	252	31	65177	65384	2233
20	04-03-	-17T22:	59:34.5	36273	38185160	.99753	12453	12771	1	2	0	262122	252	31	65166	65384	2240
20	04-03-	-17T23:	00:06.5	36284	38185192	.99753	12458	12775	1	2	0	262124	252	31	65163	65381	2241

ROSETTA	Document: Issue:	RO-IGEP-TR-0007 3			
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Date: Page:	January 20, 2010 47			

2004-03-17T23:00:38.536295 38185224.99753	12464	12779	1	2	0	262120	252	31	65165	65381	2242
2004-03-17T23:01:10.536307 38185256.99753	12470	12783	1	2	0	262123	252	31	65164	65383	2242
2004-03-17T23:01:42.536318 38185288.99753	12476	12788	1	2	0	262123	252	31	65163	65381	2243

7.2.2 Magnetic field files

FORMAT										
Filename:	RF	RPCMAG{yymmdd}T{hhmm}_RAW_{OB IB}_M{mode}.ASC								
Header	#	#								
Datacolumn	1.	Timestring (SCET in UTC)	$\{yyyy-mm-dd\}T\{hh\}:\{mm\}:\{ss\}.\{fffff\}$							
	2.	S/C CLOCK OBT	{nnnnnnnnnn}.{nnnnn}							
	3.	MAG Bx	20 bit ADC Value							
	4.	MAG By	20 bit ADC Value							
	5.	MAG Bz	20 bit ADC Value							
	6.	Sensor Temperature	20 bit ADC Value							
	7.	Quality	Value							

Example:

FILE= RPCMAG040317T2257_RAW_OB_M4.ASC # MODE= SID4 MEDIUM MODE # SENSOR= OB # SCET OBT BX[COUNTS] BY[COUNTS] BZ[COUNTS] MAGTHERM1[COUNTS] QUALITY[] 0 0 0 0 2004-03-17T22:57:28.586979 38185035.04828 -5483 -2366 35817 2004-03-17T22:57:28.786979 38185035.24828 -5490 -2372 35814 0 0 2004-03-17T22:57:28.986979 38185035.44828 -5495 -2382 35812 2004-03-17T22:57:29.186979 38185035.64828 -5497 -2394 0 0 35814 0 0 2004-03-17T22:57:29.386979 38185035.84828 -5498 -2403 35818

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 48

8 File format of PDS files

A complete set of PDS files consists of two files:

- the LABEL file *.LBL, which contains information about the data measured.
- the data file $\,{}^*\!\!.\,T\!AB$ containing an ASCII table of data.

8.1 Example: A generated PDS label file (*.LBL)

PDS_VERSION_ID	=	PDS3	
LABEL_REVISION_NOTE	=	"V1.0"	
RECORD_TYPE	=	FIXED_LENGTH	
RECORD_BYTES	=	125	
FILE_RECORDS	=	85435	
DATA_SET_ID	=	"RO-E-RPCMAG-3-EAR1-CALIBRATED-V3.	0"
DATA_SET_NAME = "ROSETTA-ORBITE	ER	EARTH RPCMAG 3 EAR1 CALIBRATED V3.	0"
PRODUCT ID	=	"RPCMAG050301T0014 CLC OB M2"	
PRODUCT CREATION TIME	=	2009-12-11T08:50:03	
PRODUCT TYPE	=	"RDR"	
MISSION TO	=	"ROSETTA"	
MISSION NAME	=	"INTERNATIONAL ROSETTA MISSION"	
MISSION PHASE NAME	=	"EARTH SWING-BY 1"	
OBSERVATION TYPE	=	"EABTH SWINGBY 1"	
INSTRUMENT HOST ID	=	"RO"	
INSTRUMENT HOST NAME	=	"BOSETTA-OBBITEB"	
INSTRUMENT ID	=	"BPCMAG"	
INSTRUMENT NAME = "ROSETTA PLAS	SM/	CONSORTIUM - FLUXGATE MAGNETOMETE	R"
INSTRUMENT TYPE	=	"MAGNETOMETER"	
INSTRUMENT MODE ID	=	"SID2"	
INSTRUMENT MODE DESC	_	"	
NORMAL MODE: 32 DRIMARY & 1 SE	-	NDARY VECTORS DER 30 SECONDS"	
TARGET NAME	=	"FADTU"	
TARGET TYPE	_	"DI ANET"	
START TIME	2	2005-03-01T00-14-48 854	
START_TIME	_	2005-03-02100.04.40.054	
NOTE - "	-	2005-03-02100.00.07.098	
MACNETIC COORDINATE SYSTEM . I	CT	TR 10000	
COODDINATE CYCTEM CENTED NAME		11F J 2000	
COURDINATE_SISTEM_CENTER_NAME	-	"DUN"	
SPACECRAFI_CLUCK_SIARI_CUUNI	-	"1/6226661.209/1"	
SPACECRAFI_CLUCK_SIUP_COUNI	-	1/66342360.09737	
START III TAN DATE VALUE	_	2452420 5102876625	
START_JOLIAN DATE VALUE	_	2453430.5102070025	
STOP_JOLIAN_DATE_VALUE	2	(141020124 71 _400E1462 27	72604 20)
SC_SON_FOSTION_VECTOR	_	(1259502.00 _200109.07	72166 07)
SC_TARGET_FUSITION_VECTOR	2	(-0.22)
SDACECRAFT ALTITUDE	-	(-3.05, 0.00,	-0.22)
CUD CDACECDAET LATITUDE	2	1391/20.337	
SUB_SPACECRAFI_LAIIIUDE	-		
NOTE	Ξ,	N/A	
NULE	=.	CUN DOCITION VECTOR	
Ine values of the keywords	50	SUN_PUSITION_VECTOR,	
SC_TARGET_PUSTITUN_VECTOR 3	inc	I SC_IARGEI_VELOCIII_VECIOR,	
are related to the ECLIPJ20) reierence irame.	
SUB_SPACECRAFI_LAIIIUDE and	1 2	SUB_SPACECRAFI_LUNGIIUDE	
are northern Latitude and e	as	stern Longitude in the standard	
planetocentric IAU_ <iarge1_< td=""><td>_IN F</td><td>ME> frame. All values are computed</td><td></td></iarge1_<>	_IN F	ME> frame. All values are computed	
for the time t= SIARI_IIME.			
Distances are given in <km></km>	, 1	velocities in <km s="">, Angles in <de< td=""><td>g>"</td></de<></km>	g>"
SPICE_FILE_NAME	=	{"ATNR_P040302093352_00091.BC",	
		"KUS_LEUUM_VO.BC",	
		"KUS_V14.TF",	
		"KUS_SA_2004_V0001.BC",	
		"KUS_SA_2005_V0001.BC",	
		"KUS_SA_2006_V0001.BC",	
		"RUS_SA_2007_V0001.BC",	

ΡΟς ΕΤΤΑ	Document:	RO-IGEP-TR-0007
$N \cup S \to I \to A$	Issue:	3
	Revision:	1
ICLD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGLF Technische Universität Braunschweig	Page:	49

		<pre>"ROS_SA_2008_V0037.BC", "ROS_SA_2009_V0048.BC", "ROS_HGA_2008_V0045.BC", "ROS_HGA_2009_V0045.BC", "ROS_BC_2V15.TI", "NAIF0009.TLS", "PCK00008.TPC", "DE403-MASSES.TPC", "DE403-MASSES.TPC", "ORER00031.BSP", "ORER00031.BSP", "ORER00091.BSP", "ORER00091.BSP", "ORER00091.BSP", "ORER00091.BSP", "ORHR00091.BSP", "ROS_RC_STRUCT_V1.BSP", "ROS_RC_STRUCT_V1.BSP", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF", "EARTHFIXEDIAU.TF",</pre>
PRODUCER_ID PRODUCER_FULL_NAME PRODUCER_INSTITUTIO	= = N NAME =	"RPC_MAG_TEAM" "INGO RICHTER" "IGEP-TU-BRAUNSCHWEIG"
DATA_QUALITY_ID	=	"N/A"
DATA_QUALITY_DESC	=	н
ONLY 'GOOD' RAW DA PROCESSING_LEVEL_ID	TA HAVE BEEN P	ROCESSED AND STORED" 3
DESCRIPTION	-	п
THIS FILE CONTAINS OUTBOARD MAGNETOME BEEN APPLIED TO TH THE S/C POSITION I	CALIBRATED MA TER ABOARD THE E RAW DATA. FI S GIVEN IN EC	GNETIC FIELD VECTOR DATA OBTAINED BY THE ROSETTA S/C. GROUND CALIBRATION RESULTS HAVE ELD IS ROTATED TO ECLIPJ2000 COORDINATES. LIPJ2000 COORDINATES AS WELL."
LBL & TAB FILE HAV	E BEEN GENERAT	" ED BY S/W: GEN_CAL_DATA, VERSION V20091209"
S/C ATTITUDE COMPU NOTE	TED USING FILE =	ATNR_FDLRMA_DAP040302093352_00091.ROS"
S/C POSITION COMPUNITE	TED USING FILE =	CORHR_FDLRMA_DA00091.ROS"
GROUND CALIBRATION NOTE	FILE: RPCMAG_ =	GND_CALIB_FSDPU_FMOB.TXT" "
INFLIGHT CALIBRATI NOTE	ON FILE: RPCMA =	G_002_CALIB_OB.TXT" "
TIMESTAMPS (UTC) C TIMESTAMPS (UTC) C IN ORDER TO CORREC FLIGHT_SOFTWARE_VER	F PRIMARY SENS F SECONDARY SE T DIGITAL FILT SION_ID =	OR VECTORS HAVE BEEN SHIFTED BY 8.20 s and NNOR VECTORS HAVE BEEN SHIFTED BY 31.95 s TER TRANSFER FUNCTION." "FIL:V1.0"
PLATFORM_OR_MOUNTIN	G_DESC =	"MAGNETOMETER_BOOM: DEPLOYED"
^TABLE	= "RPCMAG05030	1TO014_CLC_OB_M2.TAB"
OBJECT NAME	= TABLE = "RPCMAG-OB-S	ID2-CLC"
INTERCHANGE_FURMAT	= ASCII	
COLUMNS	= 9 = 9	
ROW_BYTES	= 125	
OBJECT	= COLUMN	
NAME	= "TIME_UTC"	
DATA_TYPE	= TIME - 1	
BYTES	= 1 = 26	
DESCRIPTION	= "UTC TIME OF	OBSERVATION: YYYY-MM-DDTHH:MM:SS.FFFFFF"
END_OBJECT	= COLUMN	
OBJECT	= COLUMN	
DATA TYPE	= "IIME_UBI" = ASCIT REAT	
START_BYTE	= 28	
BYTES	= 15	
DESCRIPTION	= "S/C CLOCK A	T OBSERVATION TIME, SECONDS SINCE 00:00 AT
I.I.2003: SSSSSSSS END_OBJECT	= COLUMN	
OBJECT	= COLUMN	
NAME	= "POSITION_X"	
DATA_TYPE	= ASCII_REAL	
START_BYTE	= 44	

	Document:	RO-IGEP-TR-0007
$n \cup j \in I \cup A$	Issue:	3
	Revision:	1
ICDD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGEP Technische Universität Braunschweig	Page:	50

BYTES	= 13
UNIT	= "KILOMETER"
UNIT_ID	
DESCRIPTION	= "SPACECRAFT PUSITION, X CUMPUNENT, ECLIPJ2000"
END_OBJECT	= CULUMN
OR IFCT	- COLIMN
NAME	= "POSITION Y"
DATA TYPE	= ASCIT REAL
START BYTE	= 58
BYTES	= 13
UNIT	= "KILOMETER"
UNIT ID	= "km"
DESCRIPTION	= "SPACECRAFT POSITION, Y COMPONENT, ECLIPJ2000"
END_OBJECT	= COLUMN
OBJECT	= COLUMN
NAME	= "POSITION_Z"
DATA_TYPE	= ASCII_REAL
START_BYTE	= 72
BYTES	= 13
UNIT	= "KILOMETER"
UNIT_ID	= "km"
DESCRIPTION	= "SPACECRAFT POSITION, Z COMPONENT, ECLIPJ2000"
END_OBJECT	= COLUMN
ORIECT	- COLIMN
NAME	- UDE ODU
NAME DATA TYDE	= "DA_UD" = ACCTT DEAL
CTADT DVTE	- ASCII_REAL
BVTES	- 80
INTT	- 5 - "NANOTESIA"
UNIT TD	- NANDIESLA
DESCRIPTION	- III - "MAGNETIC EIELD Y COMPONENT CALIEDATED TEMPEDATURE
CORRECTED DATA	OR SENSOR FOLID 12000"
END OBJECT	= COLUMN
HND_0D5H01	OBEONN
OBJECT	= COLUMN
NAME	= "BY OB"
DATA TYPE	= ASCTI REAL
START_BYTE	= 96
BYTES	= 9
UNIT	= "NANOTESLA"
UNIT_ID	= "nT"
DESCRIPTION	= "MAGNETIC FIELD Y COMPONENT, CALIBRATED, TEMPERATURE
CORRECTED DATA,	OB SENSOR, ECLIPJ2000"
END_OBJECT	= COLUMN
OBJECT	= COLUMN
NAME	= "BZ_OB"
DATA_TYPE	= ASCII_REAL
START_BYTE	= 106
BYTES	= 9
UNIT	= "NANOTESLA"
UNIT_ID	= "nT"
DESCRIPTION	= "MAGNETIC FIELD Z COMPONENT, CALIBRATED, TEMPERATURE
CORRECTED DATA,	OB SENSOR, ECLIPJ2000"
END_OBJECT	= COLUMN
OBIECT	- COLIMN
NAME	- COLOTIN - "OILAI ITY ELACS"
DATA TYDE	- WARTITTER
CTADT DVTE	= 116
BYTES	= 8
DESCRIPTION	= "
These flags desc	ribe the quality of the magnetic field data.
The quality is o	coded in a 8 byte string. Each character can have
the following va	ilues:
VALUE: MEAN	ING :
x prope	erty described by flag is still unknown
0 nod:	sturbance, good quality
19 spec:	fic disturbance/problems, see below
Description of t	the specific flags:
FLAG-STRING FLAG	DESCRIPTION
07054004	

FLAG-STRING FLAG DESCRIPTION
87654321
:....... x = impact not assessed
..... 0 = no disturbance
..... 1 = disturbance eliminated during data analysis
..... 2 = disturbance elimination failed
..... 3 = data disturbed

	Document:	RO-IGEP-TR-0007
$n \cup b \perp 1 \perp A$	Issue:	3
	Revision:	1
ICDD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGEP Technische Universität Braunschweig	Page:	51

:::::	:	2	IMPACT OF LANDER HEATER CURRENTS:
:::::			<pre>x = impact not assessed</pre>
:::::			0 = no disturbance
:::::			1 = disturbance eliminated during data analysis
:::::			<pre>2 = disturbance elimination failed</pre>
:::::			3 = data disturbed
:::::			
:::::		3	BUUM DEPLUYMENT:
:::::			0 = boom deployed
:::::			1 = boom stowed
:::::			2 = boom deployment ongoing. Data only valid in
:::::			instrument coordinates
			3 = pyros fired for boom release
		4	OFFERT DELATED EFFECTS.
		4	v = offset issues not assessed
			$0 = n_0 \text{ offset problems}$
			1 = offact behavior not alcon
			2 = offset drifts sonsor not in thermal
			aquilibrium thug temperature model N/A
			3 = offset drifts, reason unknown
			A = ofact jump detected reason unknown
			4 - orset jump detected, reason unknown
		5	CORRELATION RETWEEN IN AND OR SENSOR
		0	x = correlation not assessed
			0 = perfect correlation
			1 = good correlation
			2 = poor correlation
			3 = IB and OB show different long term behavior
			o ib and ob bhow difforont fong torm bonderor
:::		6	OTHER IMPACTS DECREASING THE QUALITY
::			x = no assessment
::			0 = no other problems detected
::			1 = TBD
::			2 = TBD
::			3 = TBD
::			4 = TBD
::			5 = data disturbed by AC signal originated in s/c
::			6 = data noisy due to power on failure
::			7 = data not calculatable due to thermistor failure
::			8 = sensor saturated due to huge external field
::			9 = sensor saturated, instrument power on sequence failed
::			
::		7	TBD
:			x = no assessment
:			
:		8	TBD
:			x = no assessment
:			
"			
END_OBJE	ECT		= COLUMN
END_OBJE	CT		= TABLE
END			

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 52

8.2 Example: A generated PDS data file (*. TAB)

2004-05-08T11:18:27.536951	42635892.48346	-72425206.90 -112079955.03	929862.30	-9.70	15.24	16.59 xx510003
2004-05-08T11:18:59.959398	42635925.10495	-72424287.33 -112080504.45	929863.90	-8.60	14.18	14.15 xx510003
2004-05-08T11:19:31.991236	42635957.12581	-72423378.83 -112081047.24	929865.48	-9.61	14.64	15.14 xx510003
2004-05-08T11:20:04.023074	42635989.14667	-72422470.32 -112081590.03	929867.06	-10.24	15.12	16.88 xx510003
2004-05-08T11:20:36.055019	42636021.16760	-72421561.81 -112082132.81	929868.64	-12.48	13.09	18.35 xx510003
2004-05-08T11:21:08.086872	42636053.18847	-72420653.30 -112082675.59	929870.21	-12.37	13.57	18.86 xx510003
2004-05-08T11:21:40.118710	42636085.20933	-72419744.78 -112083218.36	929871.79	-10.47	14.73	18.32 xx510003
2004-05-08T11:22:12.150654	42636117.23025	-72418836.26 -112083761.12	929873.37	-11.01	14.34	17.46 xx510003
2004-05-08T11:22:44.182492	42636149.25111	-72417927.73 -112084303.88	929874.95	-10.09	11.26	22.05 xx510003
2004-05-08T11:23:16.214330	42636181.27197	-72417019.21 -112084846.62	929876.53	-10.52	13.33	21.14 xx510003
2004-05-08T11:23:48.246275	42636213.29291	-72416110.67 -112085389.37	929878.11	-10.12	14.41	20.18 xx510003
2004-05-08T11:24:20.278113	42636245.31377	-72415202.13 -112085932.10	929879.69	-9.95	13.79	15.15 xx510003
2004-05-08T11:24:52.309951	42636277.33463	-72414293.59 -112086474.83	929881.26	-10.55	14.34	16.25 xx510003
2004-05-08T11:25:24.341911	42636309.35557	-72413385.04 -112087017.56	929882.84	-10.66	14.35	19.24 xx510003

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3	
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 53	

9 File format of LATEX Soutput file PLOTFILES.TEX

```
<u>%______</u>
PLOTS OF DAY: 2004-05-07
%
<u>%______</u>
%====== Actual Time: Thu Sep 02 09:23:16 2004
\def\fileA{\special{psfile=D:/Rosetta/rpcPds/rpcMag/work/PLOTS/CALIBRATED/2004/
                MAY/LEVEL_A/HK/RPCMAG040507T2033_CLA_HK_P2000_2400_002.EPS
                vscale = 80 hscale = 80 angle=0 }}
\begin{figure}[h]
  \begin{center}
  \setlength{\unitlength}{1cm}
  \begin{picture}(15.0,21)(0,0)
  \put(-1.0,0.0){ \fileA }
  \put( -1.00, 1.00){\framebox(17.0,21){}}
  \end{picture}
  \caption{File: RPCMAG040507T2033\_CLA\_HK\_P2000\_2400\_002}
  \label{RPCMAG040507T2033_CLA_HK_P2000_2400_002}
  \end{center}
\end{figure}
\clearpage
\def\fileA{\special{psfile=D:/Rosetta/rpcPds/rpcMag/work/PLOTS/CALIBRATED/2004/
                MAY/LEVEL_A/HK_B/RPCMAG040507T2033_CLA_HK_B_P2000_2400_002.EPS
                 vscale = 80 hscale = 80 angle=0 }}
\begin{figure}[h]
  \begin{center}
  \setlength{\unitlength}{1cm}
  \begin{picture}(15.0,21)(0,0)
  \put(-1.0,0.0){ \fileA }
  \put( -1.00, 1.00){\framebox(17.0,21){}}
  \end{picture}
  \caption{File: RPCMAG040507T2033\_CLA\_HK\_B\_P2000\_2400\_002}
  \label{RPCMAG040507T2033_CLA_HK_B_P2000_2400_002}
  \end{center}
\end{figure}
\clearpage
\def\fileA{\special{psfile=D:/Rosetta/rpcPds/rpcMag/work/PLOTS/CALIBRATED/2004/
                MAY/LEVEL_B/OB/RPCMAG040507T2339_CLB_OB_M5_T2000_2400_002.EPS
                vscale = 80 hscale = 80 angle=0 }}
\begin{figure}[h]
  \begin{center}
  \setlength{\unitlength}{1cm}
  \begin{picture}(15.0,21)(0,0)
  \put(-1.0,0.0){ \fileA }
  \put( -1.00, 1.00){\framebox(17.0,21){}}
  \end{picture}
```

ROSETTA	Document: Issue:	RO-IGEP-TR-0007		
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 54		
<pre>\caption{File: RPCMAG040507T2339_CLB_0B_M5_T2000_ \label{RPCMAG040507T2339_CLB_0B_M5_T2000_2400_002} \end{center} \end{figure} \clearpage</pre>	_2400_002}			

% DEL /Q /S D:\Rosetta\rpcPds\rpcMag\work*040507*CLA*0B*eps

% Datei wurde gelscht - D:\Rosetta\rpcPds\rpcMag\work\PLOTS\CALIBRATED\2004\MAY\ LEVEL_A\OB\RPCMAG040507T2339_CLA_OB_M5_T2000_2400_002.EPS %------

/° %-----

%-----

```
% DEL /Q /S D:\Rosetta\rpcPds\rpcMag\work\*040507*CLA*IB*eps
```

% Datei wurde gelscht - D:\Rosetta\rpcPds\rpcMag\work\PLOTS\CALIBRATED\2004\MAY\ LEVEL_A\IB\RPCMAG040507T2339_CLA_IB_M5_T2000_2400_002.EPS

%------

%-----

ΒΟςΕΤΤΛ	Document:	RO–IGEP–TR–0007
	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
ТСГГ Technische Universität Braunschweig	Page:	55

10 Ancillary IDL Routines

Additionally to the original DDS2PDS software package there are a few extra routines which ease the preparation of the archive data:

10.1 Generation of a Temperature Model

For a proper calibration of the MAG data the right temperature model is essential. It diminishes or even eliminates the temperature influence of the MAG sensor.

The sensors have been calibrated during the Ground calibration in a temperature range down to $-60^{\circ}C$. This original Temperature calibration model is applied ALWAYS to the data. It is taken into account using the two files ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_GND_CALIB_FSDPU_FMIB.TXT

ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_GND_CALIB_FSDPU_FMOB.TXT

However, during flight lower temperatures have been seen. Besides this, hysteretic effects might be possible changing the sensor characteristics in the way that a new correction model becomes necessary. For this purpose the IDL S/W

CALIB_ROS_TEMP_XXXXXX.PRO is used. For the the early mission phases CVP — EAR1 a common model (002) exists, which has been created using CALIB_ROS_TEMP_002.PRO This model simply consists of ONE 3rd order polynomial for each sensor component, derived from CVP — EAR1 data. The DDS2PDS S/W uses the results by reading the parameters from the 2 files: ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_002_CALIB_IB.TXT ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_002_CALIB_OB.TXT

It turned out that the sensor characteristics can be described much better with a model using different polynomials for shorter time intervals. So the model for the MARS and later phases is calculated on the base of daily adapted model parameters. These have been calculated with the S/W CALIB_ROS_TEMP_MARS_006.PRO. The related output files to be read by DDS2PDS are listed below:

```
ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_070223_006_CALIB_IB.TXT
ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_070223_006_CALIB_0B.TXT
```

•••

ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_070227_006_CALIB_IB.TXT ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_070227_006_CALIB_0B.TXT

The values in these files are listed in instrument coordinates. Additionally the related label files *.LBL will be generated. Only the <data_set_id> in these label files has to be adapted manually. This can easily be done by the software do_substitute_dataset_ids.pro

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 56

For the other mission phases similar files have to be generated with adapted software. In general the software CALIB_ROS_TEMP_<PHASE>_nnn.PRO generates the calibration files for the desired days: ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_yymmdd_nnn_CALIB_IB.TXT ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_yymmdd_nnn_CALIB_OB.TXT

ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_yymmdd_nnn_CALIB_IB.LBL ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_yymmdd_nnn_CALIB_OB.LBL

Here yymmdd denotes the date and nnn denotes the version of the temperature model.

ROSETTA	Document: Issue:	RO-IGEP-TR-0007
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 57

How to achieve a new temperature model?

Procedure:

• Generate GND calibrated data:

START: GEN_TEMP_MODEL= N DAY= 2007-02-23 CAL_FILE_ID= GND GEN_LEV_E_DATA= Y RESAMPLED_AVERAGE= 200 BOOM= DEPLOYED GET_DATA= N CAL_DATA= Y GEN_PLOTS= N ACTION: END:

• Generate LEVEL_A and LEVEL_B GND plots:

```
START:
REFERENCE_BODY= MARS
CAL_FILE_ID= GND
DAY= 2007-02-23
BOOM= DEPLOYED
GET_DATA= N
CAL_DATA= N
GEN_PLOTS= Y
GEN_RAW_PLOTS= Y
GEN_LEV_A_PLOTS= Y
GEN_LEV_B_PLOTS= Y
GEN_DYNSPEC_PLOTS= N
GEN_POW_SPEC_PLOTS= N
SPEC_SOURCE= LEVEL_A
PLOT_TIME_START= 00:00:00
PLOT_TIME_END= 24:00:00
ACTION:
END:
```

• Glue the generated plots together in a common document for an easy scrolling and inspection:

	Document:	RO–IGEP–TR–0007
$\Pi \cup \Im \Box \Box \Box \Box A$	Issue:	3
	Revision:	1
ICED Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
IGET Technische Universität Braunschweig	Page:	58

Use: glue_filenames_eps.pro Change directory: cd \ROSETTA\rpcPds\rpcMag\Tmp\ Execute: glue_eps.bat This procedure will produce a plot-overview document in rpcmag\plots\overview\<timeperiod>.ps from the desired GND plots located in the rpcmag\plots\calibrated\...\LEVEL_B\OB directory. For an easy access of the plots (page numbering!) the overview document should be converted to PDF.

• Tempmodel Generation: Generate averaged (A200), GND calibrated LEVEL_E data for the desired time period. This is done with the GEN_TEMP_MODEL=Y option in the following way: (example) :

```
START:
GEN_TEMP_MODEL= Y
DAY= 2007-02-23
CAL_FILE_ID= GND
GEN_LEV_E_DATA= Y
RESAMPLED_AVERAGE= 200
BOOM= DEPLOYED
GET_DATA= N
CAL_DATA= Y
GEN_PLOTS= N
ACTION:
END:
```

The generated data will be written to D:\ROSETTA\DATA\CALIB\TEMPMODEL_AVG.

- Move the just produced data to the actual (possibly new) folder D:\ROSETTA\DATA\CALIB\TEMPMODEL_nnn where *nnn* denotes the number of the new temperature model.
- Ensure the existence of the the polynomial coefficients files ROSETTA\DATA\CALIB\P_COEFF_IB.TXT ROSETTA\DATA\CALIB\P_COEFF_0B.TXT

At least dummy files containing the start and end lines like

2000-01-01T00:00:00	0.0	0.0	0.0	0.0	0.0	0.0
2000-01-01T00:00:00	0.0	0.0	0.0	0.0	0.0	0.0
2000-01-01T00:00:00	0.0	0.0	0.0	0.0	0.0	0.0
2024-01-01T00:00:00	0.0	0.0	0.0	0.0	0.0	0.0
2024-01-01T00:00:00	0.0	0.0	0.0	0.0	0.0	0.0
2024-01-01T00:00:00	0.0	0.0	0.0	0.0	0.0	0.0

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 59

should be present. (ATTENTION: There shouldn't be any entries for days which have not been processed in the right way!) These files will be used to store and read the polynomial coefficients for each day. For every day three lines are used for the X,Y and Z component. The columns are related to TIME, a0 ... a5, the nth order polynomial coefficients.

The coefficients have to be stored in files, because the calibration routine for day (n) looks for the parameters of day (n-1) to ensure that there is no offset jump at the time of the file change. It means that the constraint $P_{n+1}(T_n(0)) \stackrel{!}{=} P_n(T_n(0))$ is fulfilled.

- In the S/W CALIB_ROS_TEMP_<phase>_nnn.PRO the desired days have to be adapted. Here also time intervals of data to be used and time intervals not to be used can be chosen for every day.
- After the successful run the following files will be available:
 - The needed calibration files (coefficients in Instrument coords.)
 ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_<id>_CALIB_IB.TXT
 ROSETTA\rpcPds\rpcMag\CALIB\RPCMAG_<id>_CALIB_OB.TXT
 - Plots of the uncorrected data (s/c coords.)
 ROSETTA\DATA\CALIB\RPCMAG_<id>_IB_T_A200_SC_COORDS_ORIGnnn.EPS
 ROSETTA\DATA\CALIB\RPCMAG_<id>_OB_T_A200_SC_COORDS_ORIGnnn.EPS
 - Plots of the corrected data (s/c coords.)
 ROSETTA\DATA\CALIB\RPCMAG_<id>_IB_T_A200_SC_COORDS_FITTEDnnn.EPS
 ROSETTA\DATA\CALIB\RPCMAG_<id>_OB_T_A200_SC_COORDS_FITTEDnnn.EPS
 - Plots of the Fit quality (s/c coords.)
 ROSETTA\DATA\CALIB\RPCMAG_<id>_IB_T_A200_SC_COORDS_TESTnnn.EPS
 ROSETTA\DATA\CALIB\RPCMAG_<id>_OB_T_A200_SC_COORDS_TESTnnn.EPS

Unused time intervals for the model creation are shown in red color.

- The offsets of the data can be tuned by setting the vectors OFF_OB and OFF_IB in CALIB_ROS_TEMP_<phase>_nnn.PRO to appropriate values. Positive values of these variables will shift the time series up, negative will shift them down.
- These procedures have to be applied iteratively until a reasonable result has been achieved.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 60

10.2 Generation of REACTION WHEEL Plots

For the assessment whether the data are disturbed by ROSETTA'S Reaction wheels the actual frequencies of the Reaction wheels have to be known and to be compared with the dynamic spectra of the magnetic field data.

There are two possibilities to get the RW data:

- Get the data directly from the DDS and convert it manually (refer to section 10.2.1). This is the old method.
- Get the data from the Imperial College server, already in ASCII format (refer to section 10.2.2). This is the standard procedure.

10.2.1 Reaction Wheel data from DDS-Server

Procedure:

- Retrieve RW-parameters from the ROSETTA DDS located at http://rodds.esa.int/
 - go Back to main page
 - Build Request
 - Request Catalog TLM
 - Data Source 196
 - On Screen \rightsquigarrow Catalog Browsing. Select 3 25 27 DATA \rightsquigarrow selection Window appears
 - select desired times GTE / LTE
 - Start transfer via ftp,
 - Type in convenient filename, leave Directory as it is, ftp target home, as it is \sim data will be sent to xanthus.sp.ph.ic.ac.uk /rosetta/rpcData/rpcTmp
- Save this TM file in d:/ROSETTA/DATA/REACTION_WHEELS/TM_DATA (*.DAT)
- File contains the following Reaction wheel parameters:

Parameter	APID	SPID	Byte_Offset	$Bit_{-}offset$	Type	Subtype	P1Val
NAAD 6014	196	4533	210	3	3	25	27
NAAD 6024	196	4533	212	3	3	25	27
NAAD 6034	196	4533	214	3	3	25	27
NAAD 6044	196	4533	216	3	3	25	27

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	1 January 20, 2010 61

• additional conversion information:

PTC	PFC	PAdded		
3	9	13	Ν	13bits unsigned Integer
				Calibration ID: 478

- Calibration Factor from mcf.dat: 0.50813
- START RVT IDL SOFTWARE:
 - @rvt in d:/rosetta/rpcpds/rpcmag/software/pcwin/idl/tlm_ascii/rvt
 - LOAD_DESCRIPTION: ./DES/ros_Wheels.des
 - OPEN TM file in d:/ROSETTA/DATA/REACTION_WHEELS/TM_DATA
 - PLOT: SAVE_TO_ASCII FILE
 - SELECT : NAAD 6014 6024 6034 6044

10.2.2 Reaction Wheel data from IC-Server

On the IGEP-FTP server an cron job is running, which executes a RSYNC command two times a day. These procedures synchronize the ftp-server with the Imperial College server XANTHUS, where also the reaction wheel data are stored in the schk7.txt file on a daily basis. This file will automatically be read during the GET_DATA process and copied to the local directory

d:\rosetta\data\reaction_wheels\ascii_data\schk7_yyyy_mm_dd.txt If a filename of this structure (i.e. SCHK is part of the filename!) is stated as RW_FILE= in the command file, the system will find the right RW data for the correction needed.

10.2.3 Processing of the Reaction Wheel data

The Reaction wheel data will be taken into account during the GEN_LEVEL_H_DATA process.

Additionally it is possible to generate plots via the *reaction_wheels.pro* routine.Here the actual settings have to be adapted in source code! Plots can then be found in it D:/ROSETTA/DATA/REACTION_WHEELS/PLOTS

ΡΟςΕΤΤΛ	Document:	RO–IGEP–TR–0007
$\mathbf{N} \cup \mathbf{O} \to \mathbf{L} + \mathbf{I} + \mathbf{A}$	Issue:	3
	Revision:	1
ICTD Institut für Geophysik u. extraterr. Physik	Date:	January 20, 2010
Ι GΓΓ Technische Universität Braunschweig	Page:	62

10.3 Generation of Solar Array (SA) and Hight Gain Antenna (HGA) Plots

Plots of the Solar Array angles and the High Gain Antenna Orientation can be generated by the routine

calc_rosetta_hga_solar_arrays.pro

Before running the procedure the entries have to be changed according to the desired processing day.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 63

10.4 Generation of overview plots and overview about available data

To get an idea at which days data are available in which calibration level several routines exist to generate a complete report. The generation process is the following:

- Generate a list of all available data using create_list_rpcmag_data_overview.pro Check actual parameters in source code before running. This procedure generates the following list: D:/ROSETTA/DATA/AVAILABILITY/RPCMAG_OVERVIEW.LST (Name can be adapted). This item takes a long time
- Using this list overview plots about available data level for each day can be produced. START: create_plot_rpcmag_data_overview.pro (quick step) Check actual parameters in source code before running. This routine generates *.PS and *.TEX file for the generation of an overview report. Output is written to D:/ROSETTA/DATA/AVAILABILITY
- 3. An automatic quality check of all LEVEL. F data can be performed using the routine $quality_check.pro$

This test is based on checking the thermal equilibrium of the sensors. If a certain threshold of the change rate of the sensor temperature differences is exceeded, the data will be marked as bad. Output is

- Text files with the time intervals of bad data: D:/ROSETTA/DATA/QUALITY/<datasetID>/RESAMPLED/yyyy/month/RPCMAGyymmdd_CLF_QUALITY_FLAGS.TXT
- Plots of LEVEL_F data with red tagged bad intervals: D:/ROSETTA/DATA/QUALITY/<datasetID>/RESAMPLED/yyyy/month/RPCMAGyymmdd_CLF_OBIB.PS
- Plots of differences (OB-IB) of LEVEL_F data with red tagged bad intervals: D:/ROSETTA/DATA/QUALITY/<datasetID>/RESAMPLED/yyyy/month/RPCMAGyymmdd_CLF_DIFF.PS
- A file named D:/ROSETTA/DATA/QUALITY/PLOTFILES.TEX which contain a LAT_FX compliant input file of all plots and figures for the report.

Adapt the parameter of the routine inside the source code before running.

4. Now the report can be generated. Use Template for RO-IGEP-TR0017 to produce the desired document.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010

10.5 Generation of a Operations Logbook

The PSA/PDS requested an operation logbook included in the data archive. Such a logbook can easily be created with the IDL S/W $GEN_LOGBOOK.PRO$.

The input section has to be adapted for each mission phase. The software requires a START_DATE and an END_DATE for the time interval to be analyzed and logged. Additionally a flag has to be set which indicates whether the TC logging files $(rpc\{yymmdd\}dds00_cmh_cmdpf.txt)$ and event files (events.txt) created on the ROSETTA DDS have still to be retrieved from the IGEP ftp server or are already available on the local machine. These files contain the information which will be extracted by the $GEN_LOGBOOK$ Software.

Basic principle of the Software:

- The software compiles a list of all dates where RPC has been switched on in the stated time interval between START_DATE and END_DATE. This information is derived from the magnetic field raw data filenames present the *D:/ROSETTA/DATA/ASS_FILES* directory (i.e a logfile can not be created until the magnetic field data have been treated for the desired days.)
- If the GET_DATA flag is set to 1 , the s/w transfers all TM- and EVENTS- files needed from the IGEP ftp-server to the local machine and stores them in the D:/ROSETTA/DATA/TLM directory.
- Now the logfile with the name LOGBOOK_{Startdate}_{Enddate}.ASC and the related PDS LABEL-file LOGBOOK_{Startdate}_{Enddate}.LBL will be created in the D:/ROSETTA/DATA/TLM directory.
- The logfile will then contain chronological information about all TCs sent to RPC-MAG, all WARNINGS and special EVENTS affecting the instrument. All this info is automatically extracted from the mentioned DDS logging files
- After the automatic creation further information can be added manually. The manually created lines should be assigned by a preceding ";" to be able to distinguish between manually and automatically created lines which start with a "#".
- Finally the Logbook files (*.*TXT*, *.*LBL*) have to be moved to the actual DOCUMENT directory. Also the actual *DOCINFO.TXT* has to adapted.

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 65

11 Archive Generation Procedure

This section is a brief do-list for the generation of a data set.

- 1. Generate the right *PDS.CMD* file using individual sub-files (FILE= $_{\sqcup}$) for each of the following tasks:
 - Get TM data from the ftp server and convert it to ASCII data (GET_DATA=_Y).
 - GND calibrated data have to be created for development of the actual TEMP_MODEL (refer to 10.1).
 - The temperature model has to be created using CALIB_ROS_TEMP_XXXXX.PRO
 - RAW and CALIBRATED data have to be created (CAL_DATA=_Y).
 - Time series plots and spectra have to created.
 - If any Lander heater influence is visible, corrected LEVEL_B data tables have to be created (manually or with convenient s/w). The corrected tables have to be written to D:/ROSETTA/DATA/HEATER. LEVEL_K and LEVEL_L data can be generated from this source.
 - To get an idea about the Reaction wheel influence the Reaction wheel investigation has to be done (refer to section 10.2)
 - Now all needed input for any stage of resampled LEVEL_E, LEVEL_F, LEVEL_G, LEVEL_H, LEVEL_I data is available.
- 2. Create a data quality overview and availability report. Refer to section 10.4
- 3. After generation of all data the quality dummies have to be replaced by real Quality flags. For this task a convenient quality index file D:/ROSETTA/DATA/QUALITY/QUALITY_<missionphase>.TXT has to be available. This file should be generated manually from a template (*.XLS) by optical analysis of the overview plots in the data-overview-document (TN0017, TN0022,...). This file must be saved also as a MS-DOS *.TXT file. Start the replace sequence with SET QUAL FLACS= X at DDS2PDS

Start the replace sequence with $_SET_QUAL_FLAGS=_{}Y$ at DDS2PDS.

- 4. Copy the coarse structure from former datasets.
- 5. Create a logbook for the actual missionphase. Refer to section 10.5 for details.
- 6. The data_set_id entries in the CALIB and DOCUMENT files have to be changed. This can easily be done automatically using the s/w do_substitute_dataset_ids.pro. For this task the files have to be readable and writable (remove -r flag).

Also the $PRODUCT_CREATION_TIME$ entries will be updated automatically to the actual time.

О Ѕ Е Т Т А	Document: Issue:	RO-IGEP-TR-0007 3	
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 66	

7. After generation of all ancillary files (DOCUMENT, CALIB,...) in the actual PSA directory the line width of the new *.LBL files has to be set to 80 characters (78 + CR LF). This can be done with the Software do_format_psa_file.pro . A log file is written to ../TMP/LBL_CHECK.LOG. In this file all lines are marked which exceed the nominal line length. This software also checks the *.CAT files, which have to have a line length of 72 chars (70 + CR LF). An additional log file is written to CAT_CHECK.LOG which indicates the catalog lines being to long.

Last but no least all auxiliary files

- AAREADME.TXT
- $\bullet \ CALINFO.TXT$
- CATINFO.TXT
- DOCINFO.TXT
- INDXINFO.TXT

are checked for correct line lengths as well. They have to consist of lines with 80 chars (78+ CR LF).

All document ASCII text files need the extension (*.ASC) and not (*.TXT). The line length has to be 80 chars (78 + CR LF).

$D \cap C F T T \Lambda$ Document: RO-IGEP-TR	L-0007
$\Pi \cup \cup \Box \perp \perp \perp \Pi$ Issue:	3
Revision:	1
ICED Institut für Geophysik u. extraterr. Physik Date: January 20	, 2010
IGLГ Technische Universität Braunschweig Page:	67

12 Notes

- The $DATA_SET_TERSE_DESC$ entry allows only 255 chars
- LINE CHECK:

A check of illegal chars and correct CR/LF setting can be done using the LINE tool. The is useful for the *.TXT and *.CAT files.

```
c:\programme\pds_tools\bin\line -of D:\report.txt -if <inputfile>
```

- A PDS Label Verifier is available using c:\programme\pds_tools\bin\lvtool -f <file.lbl> -r D:\report.txt -d C:\programme\pds_tools\pdsdd.idx
- Deleting files , with prompt and subdirectories: del /P /S <file>

ROSETTA	Document: Issue:	RO–IGEP–TR–0007 3	
IGEP Institut für Geophysik u. extraterr. Physik Technische Universität Braunschweig	Revision: Date: Page:	January 20, 2010 68	

12.1 Plotting Parameters

The form of appearance of the dynamic spectra strongly depends on the right choice of plot parameters. The following list gives recommendations of reasonable values for specific spectra :

SENSOR-MODE	PRIMSID2	PRIMSID3	SECSID3	SECSID3	PRIMSID4
Remark			short times	long times	
DYNSPEC_WIN_WIDTH	8	11	8	11	11
SMOOTHING_WIN	0		0(5)	0(5)	0
PLOT_AMP_MIN	1e-4	1e-4	1e-4 (1e-3)	1e-4 (1e-3)	1e-4
PLOT_AMP_MAX	10	1	10	10	1